

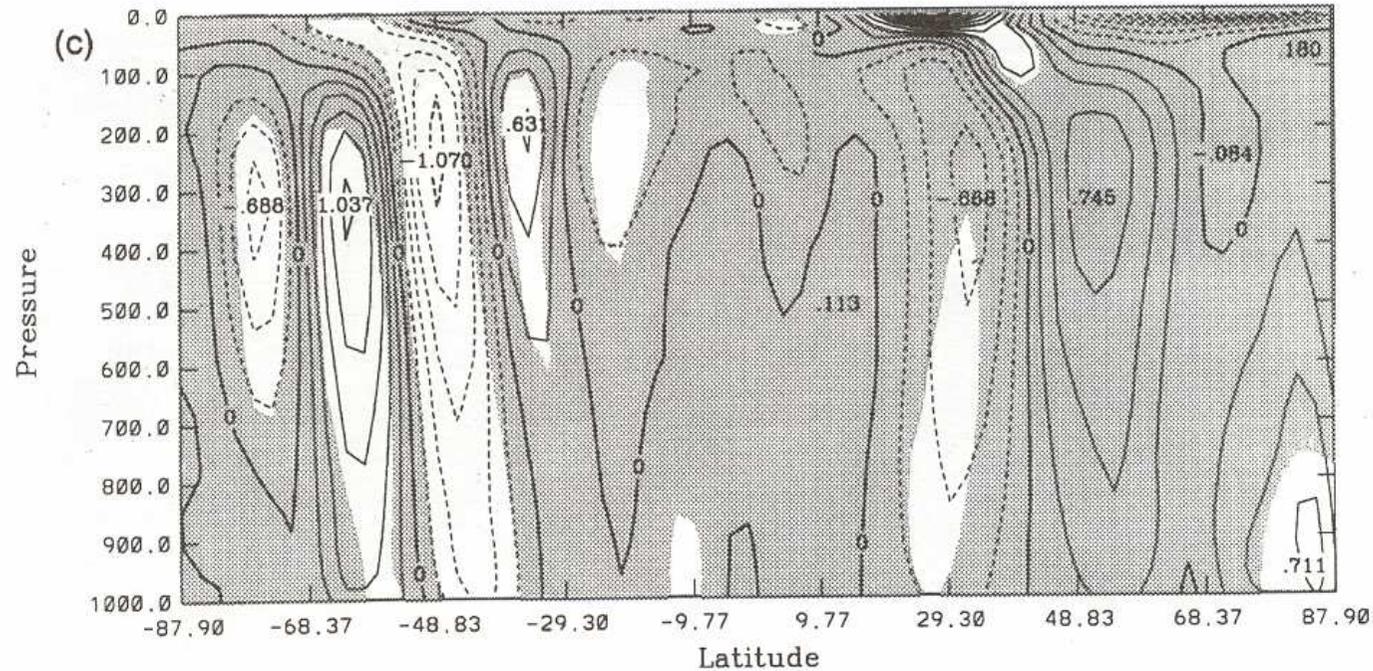
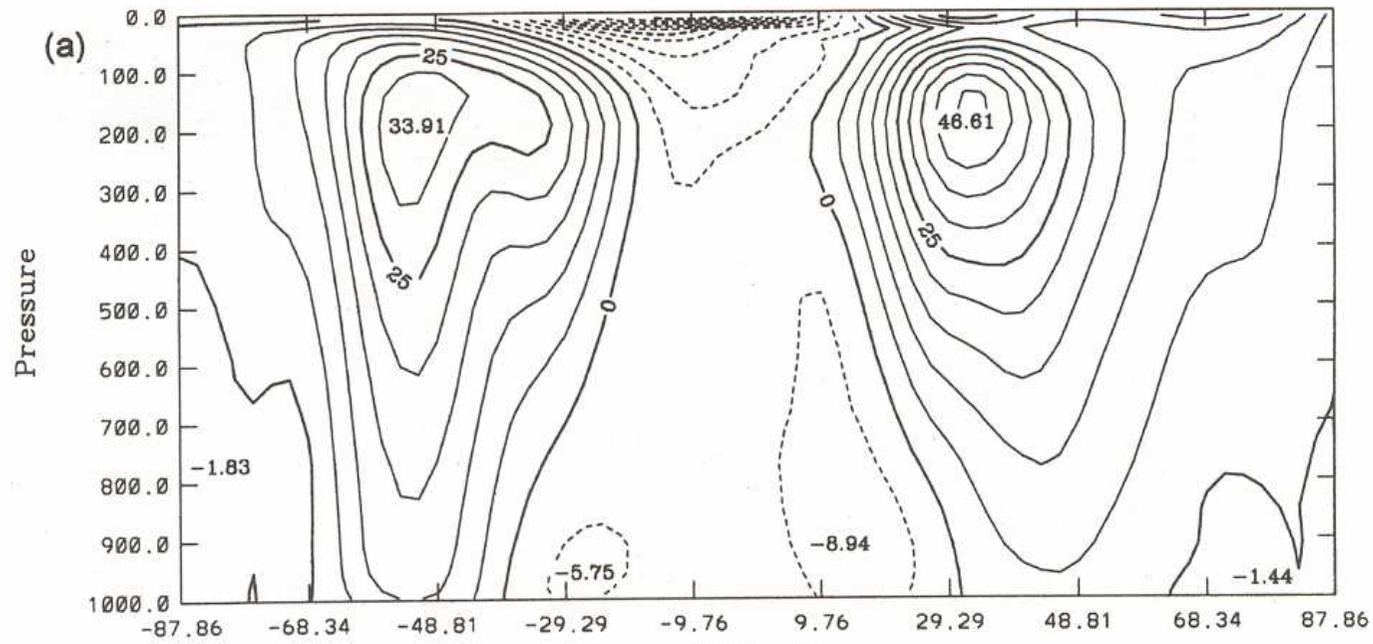
The response of tropospheric circulation to perturbations in lower stratospheric heating

Joanna Haigh Imperial College London
Mike Blackburn University of Reading
Rebecca Day Imperial College London

Outline

- GGM studies of response to solar UV.
- Sensitivity experiments.
- Comparison of model results with observational data.
- Understanding the response using a “dynamical core” model.

UGCM zonal wind



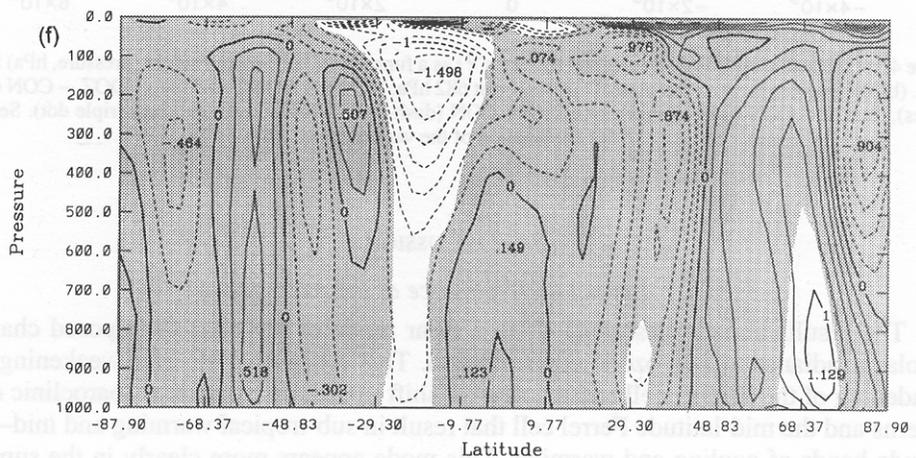
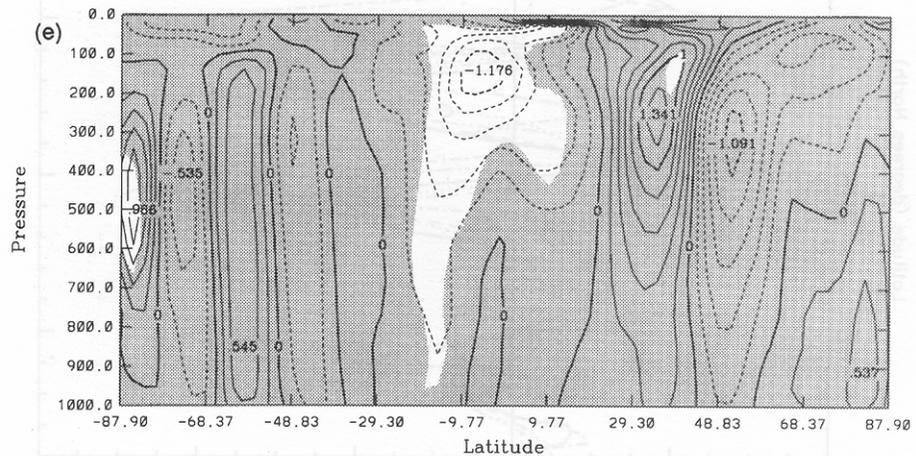
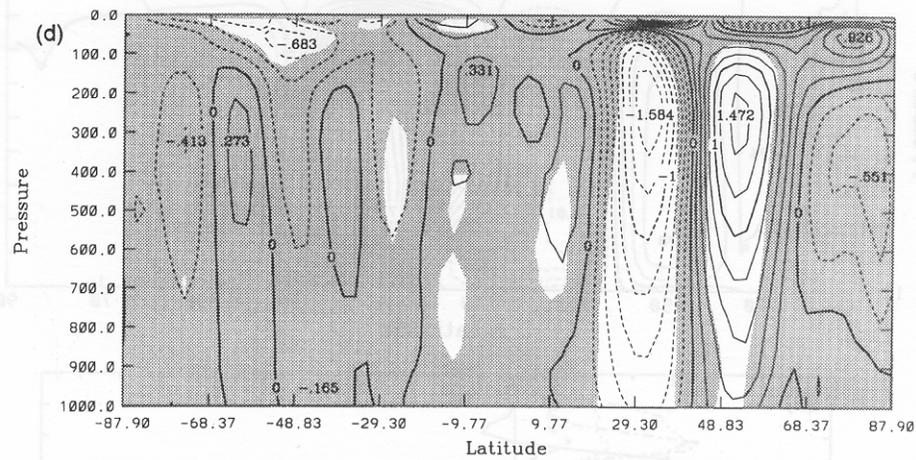
solarmax-solarmin

(2D model ΔO_3)

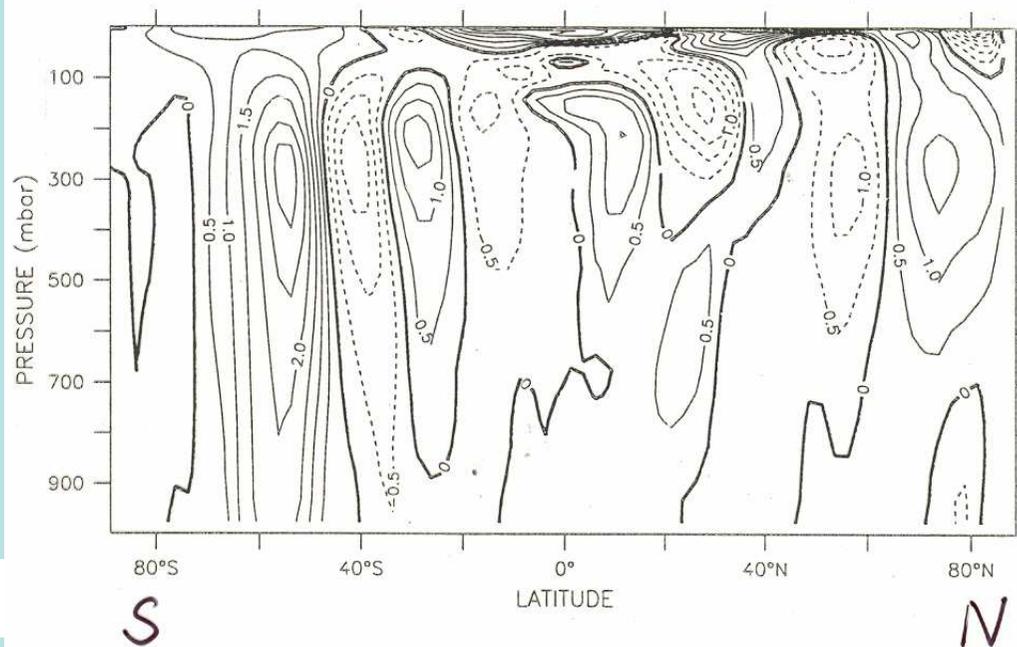
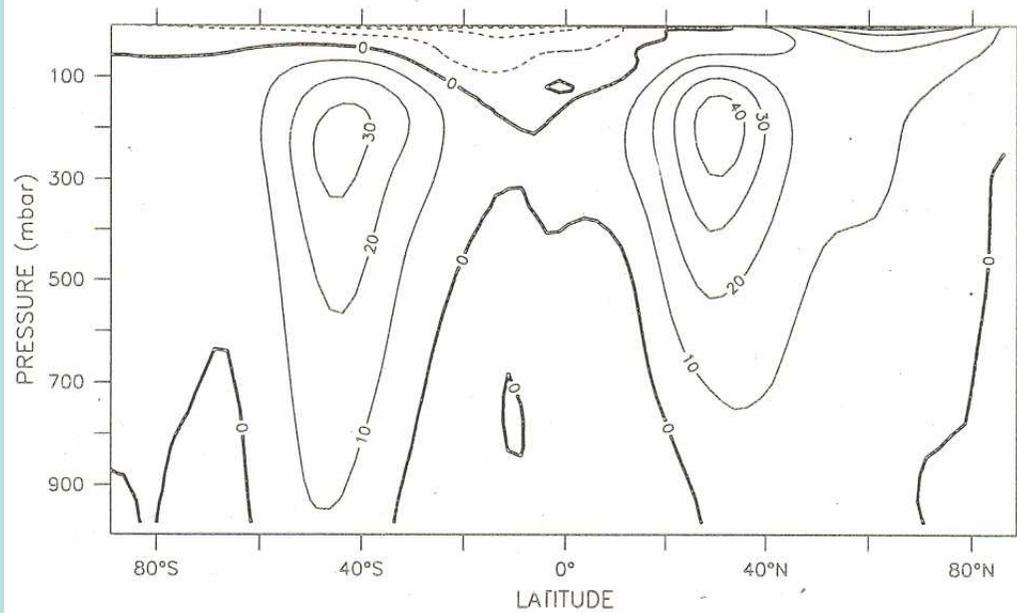
Haigh
(Science1996;
QJRMS1999)

Further UGCM solarmax- solarmin zonal wind results

Haigh (1999)

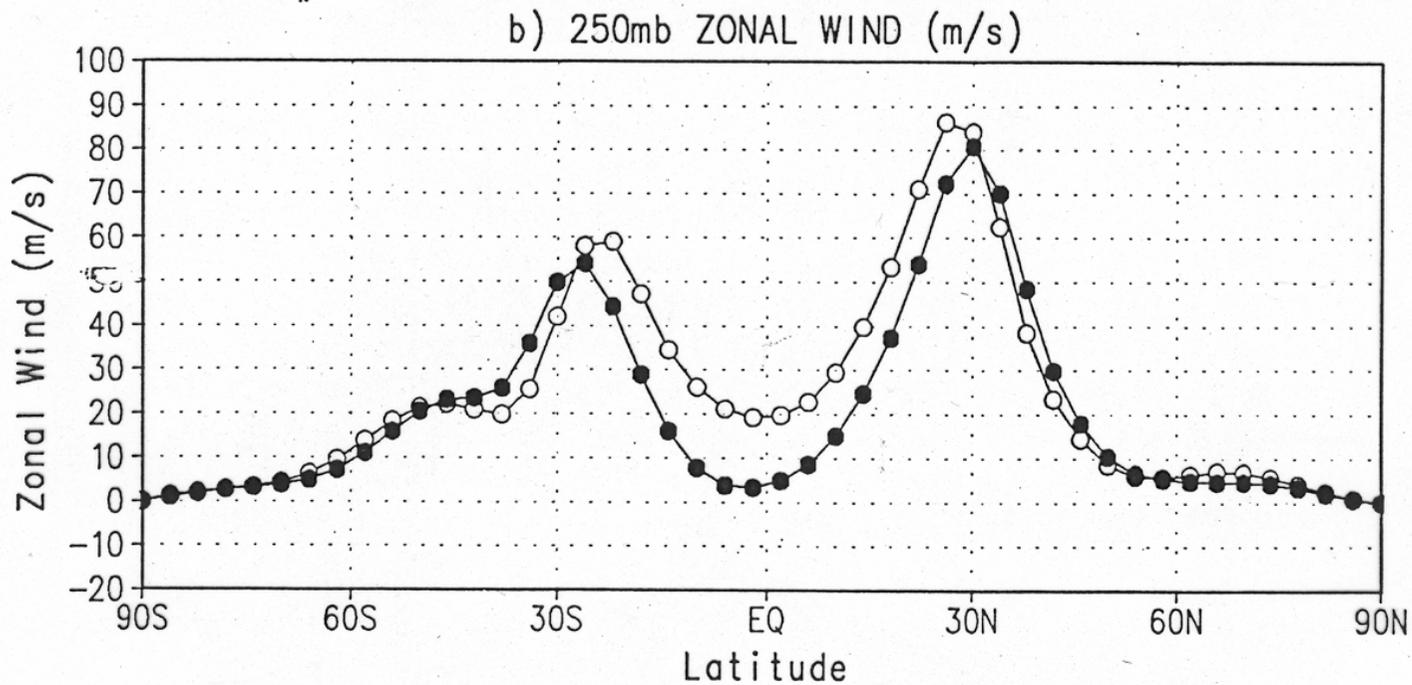
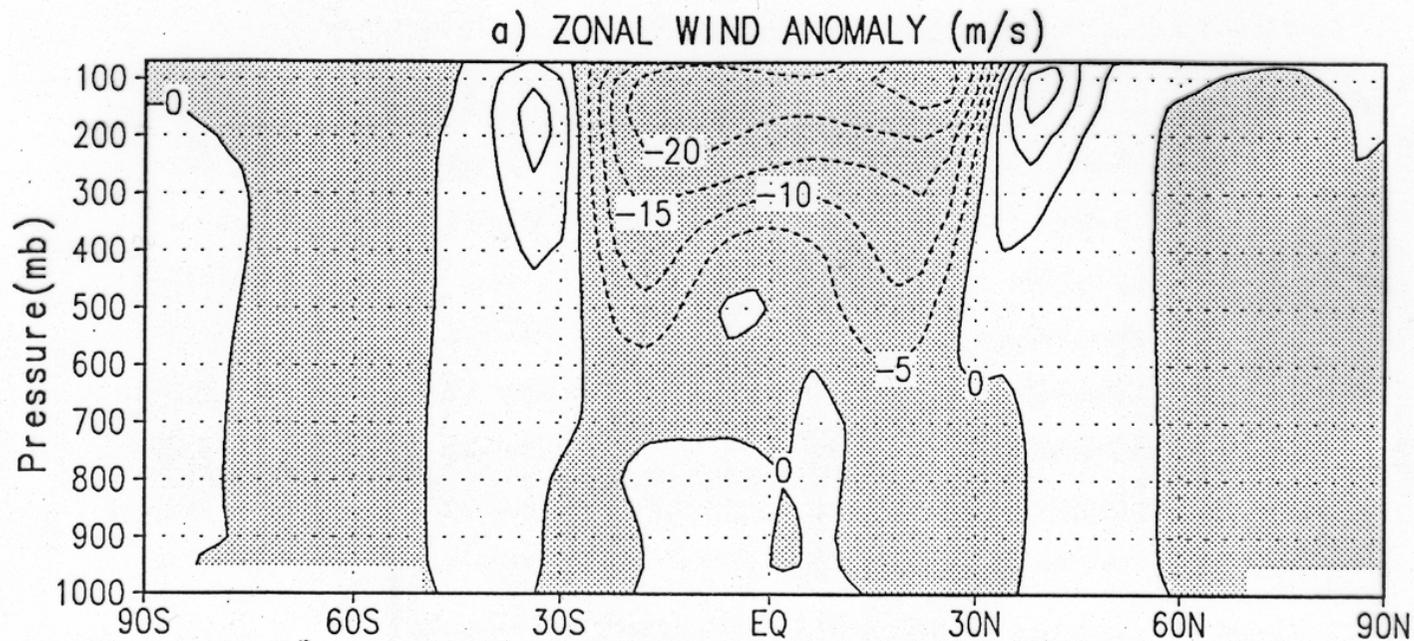


UM zonal wind



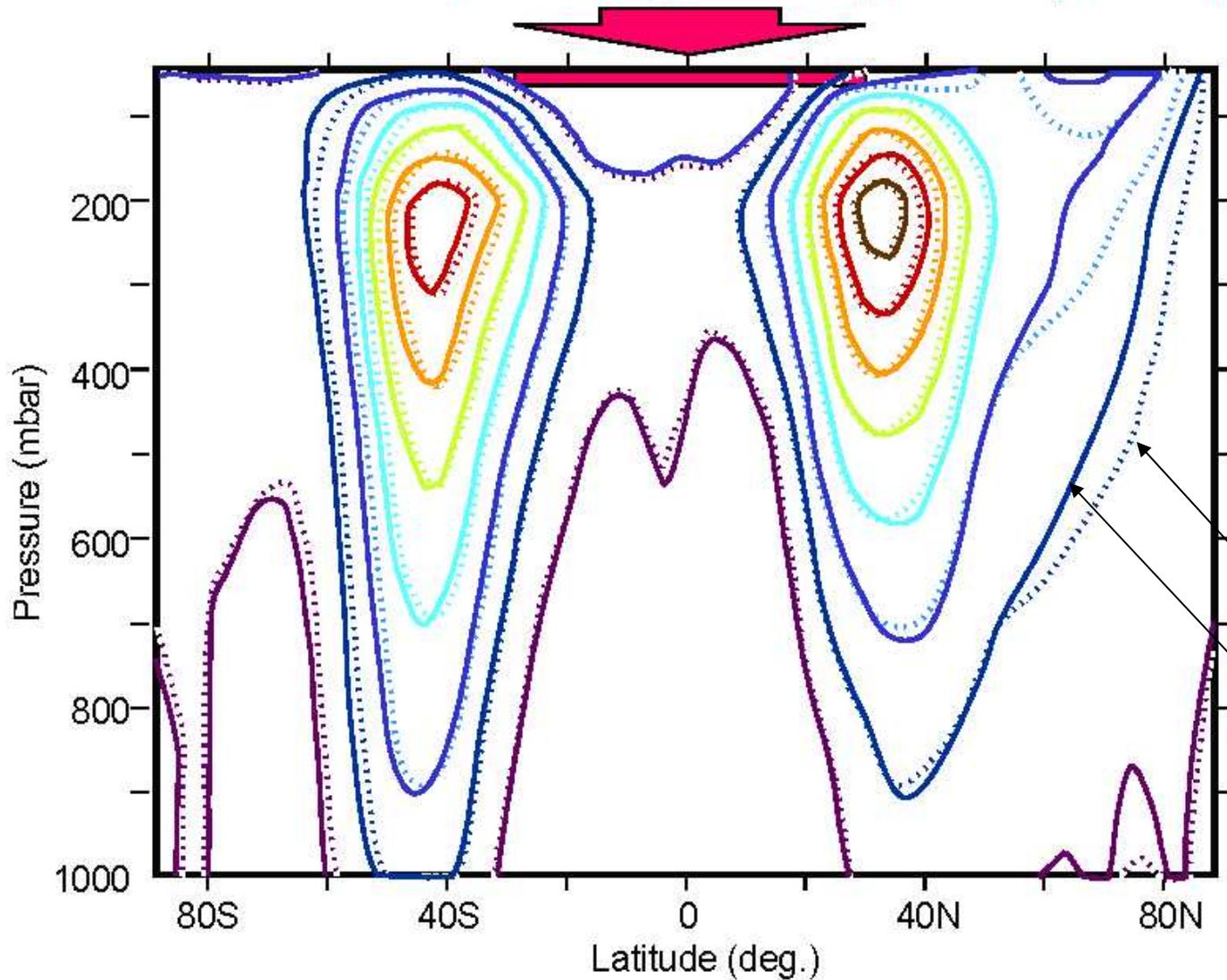
solarmax-solarmin

Larkin et al (2000)



Hou (2000)
GCM
response to
“artificial”
 δO_3

$dT/dt = 0.02 \text{ K day}^{-1}$ in daylight hours (power = $c_p \rho dT/dt$)



Westerly component of wind U_w (m s^{-1})



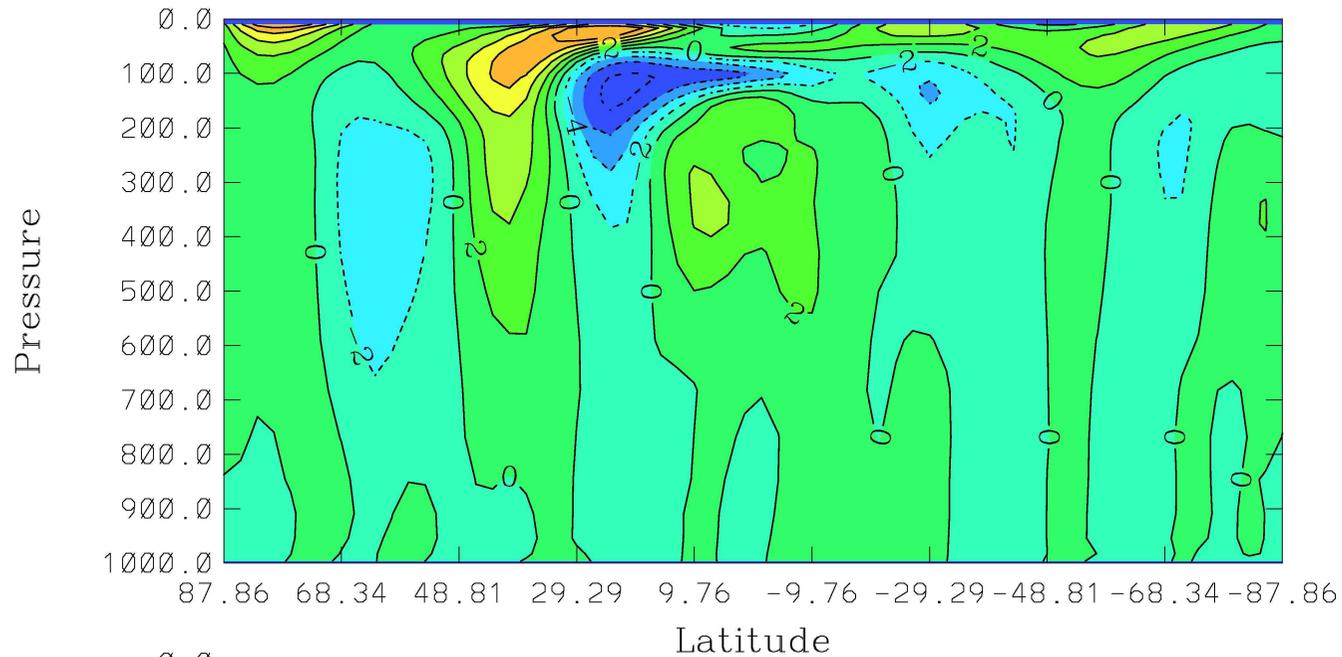
UM zonal
wind
response to
heating of
tropical lower
stratosphere

without heating

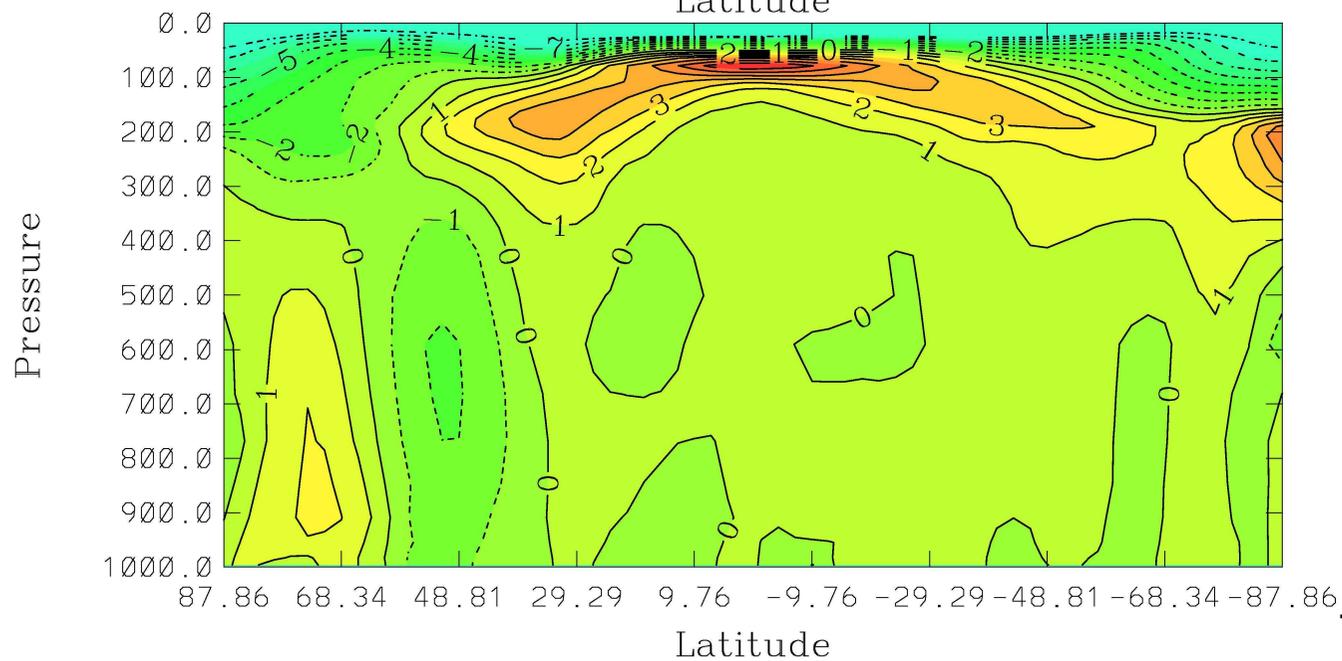
with heating

UGCM O₃ downward shift experiment

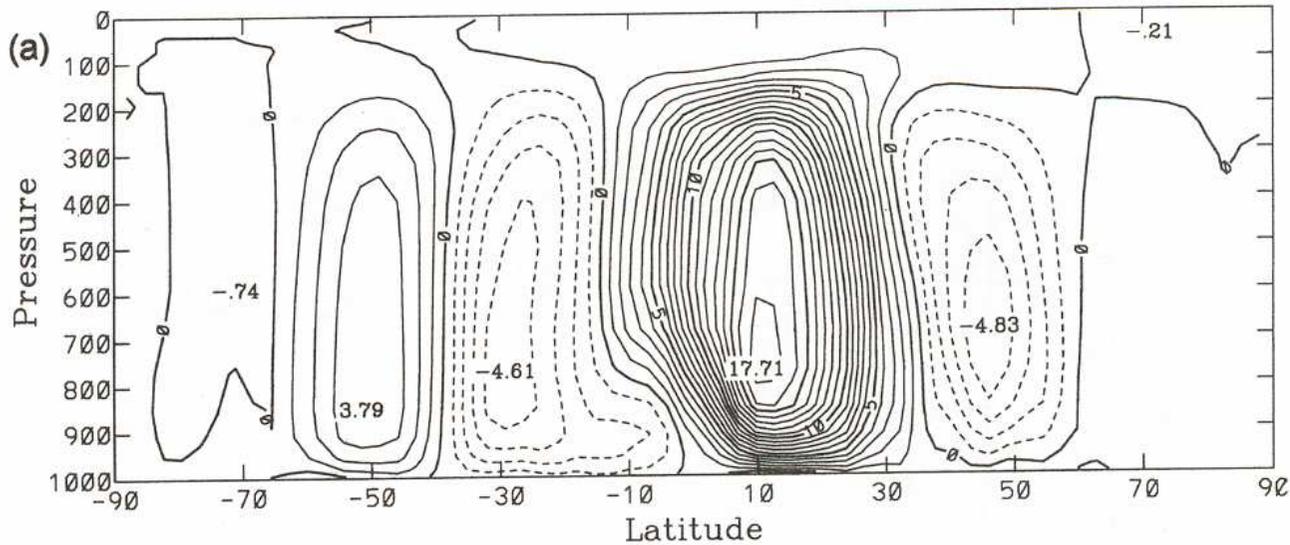
Δu (m s⁻¹)



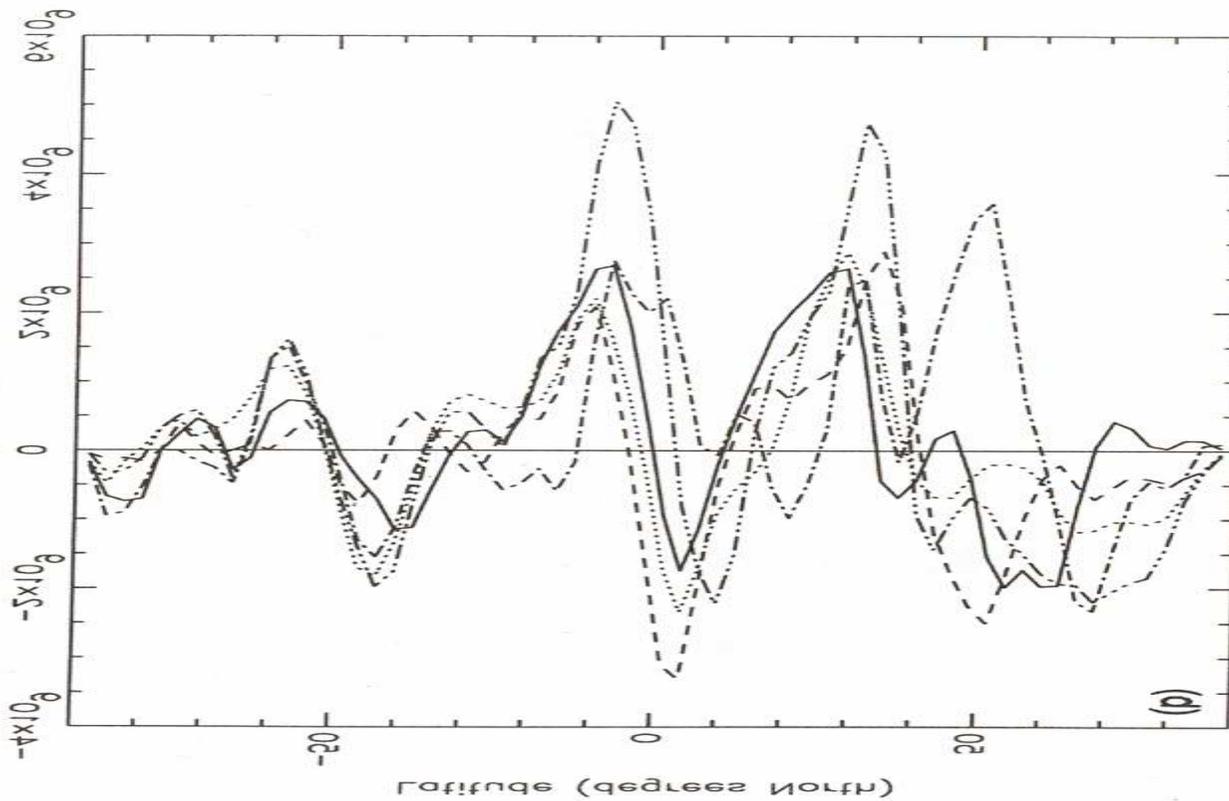
ΔT (K)



Thuburn & Craig (2000)



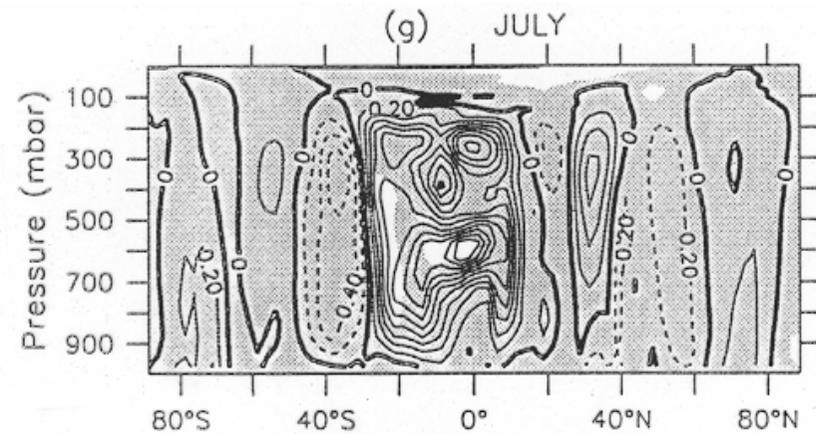
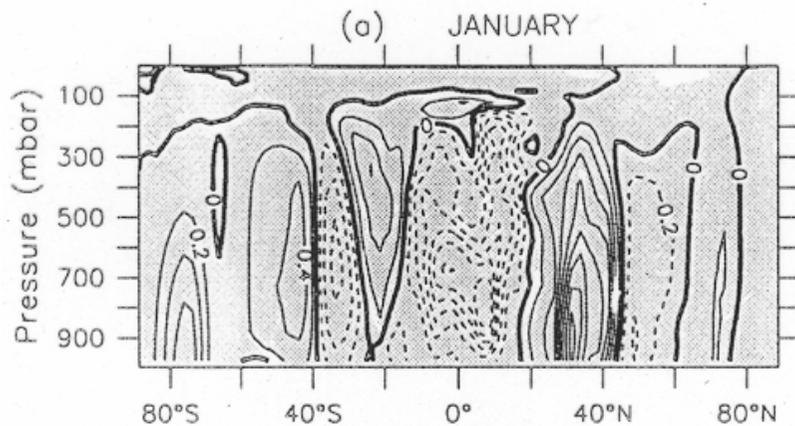
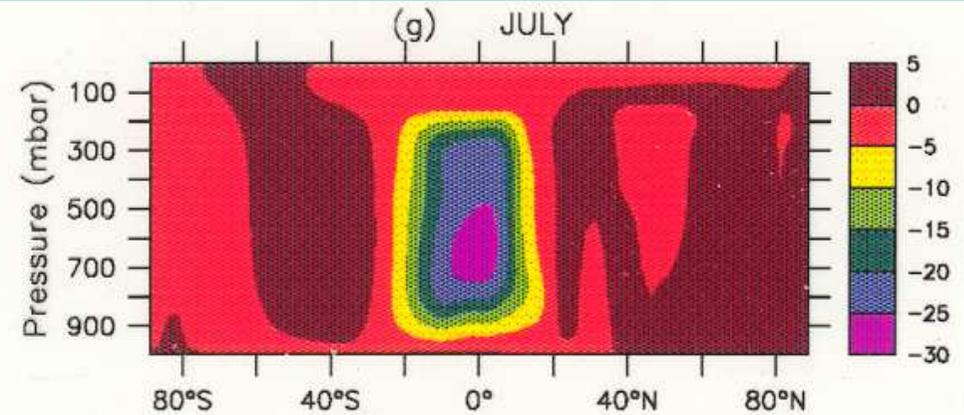
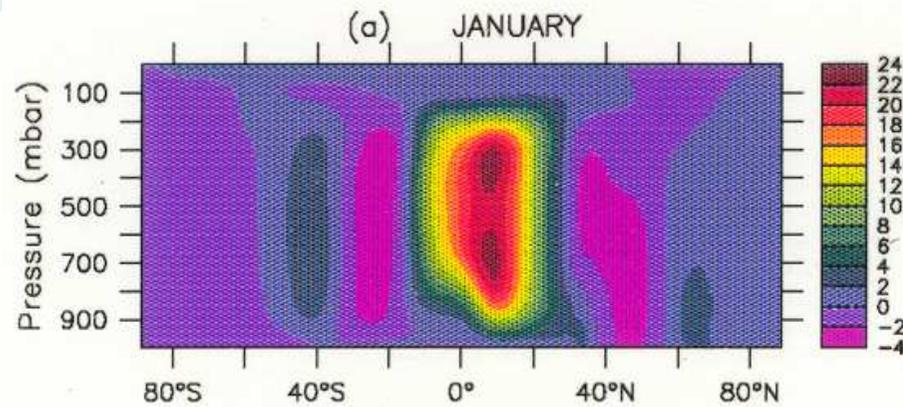
UGCM mean meridional circulation



change in MMC
at 682hPa

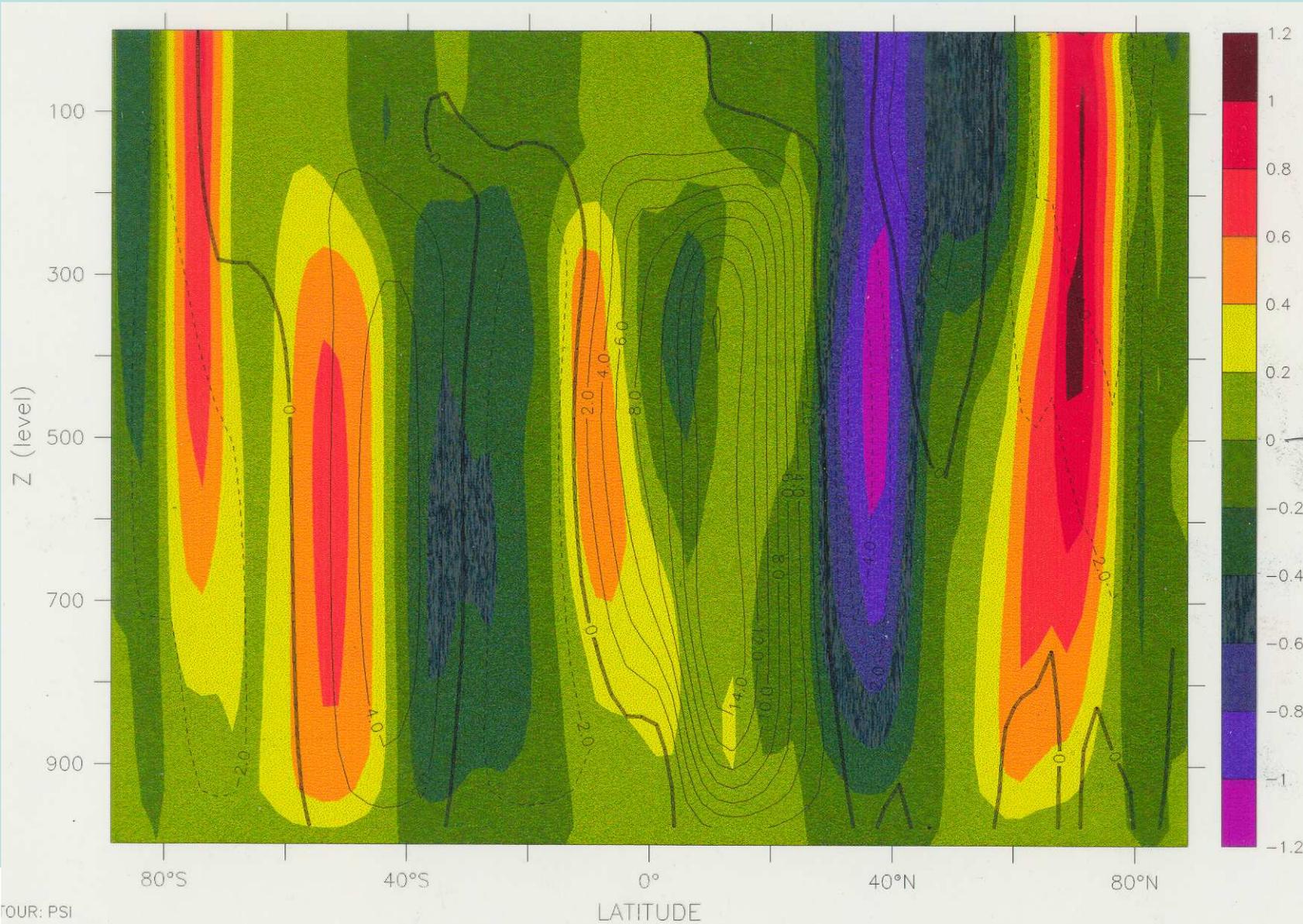
Haigh (1999)

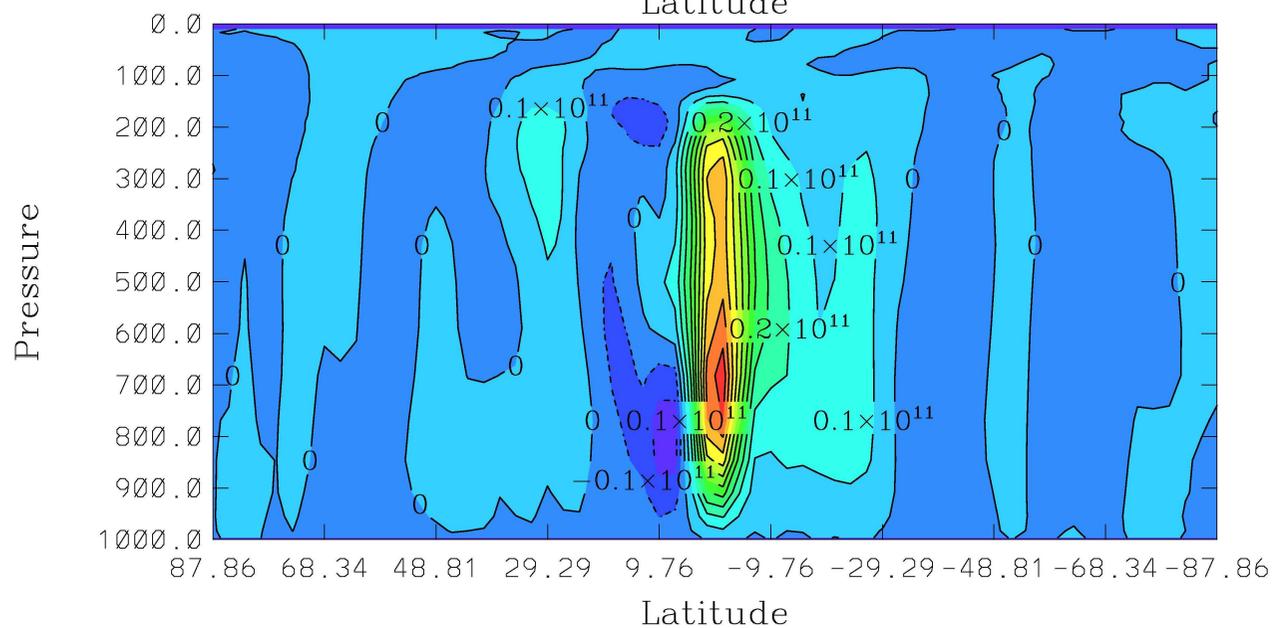
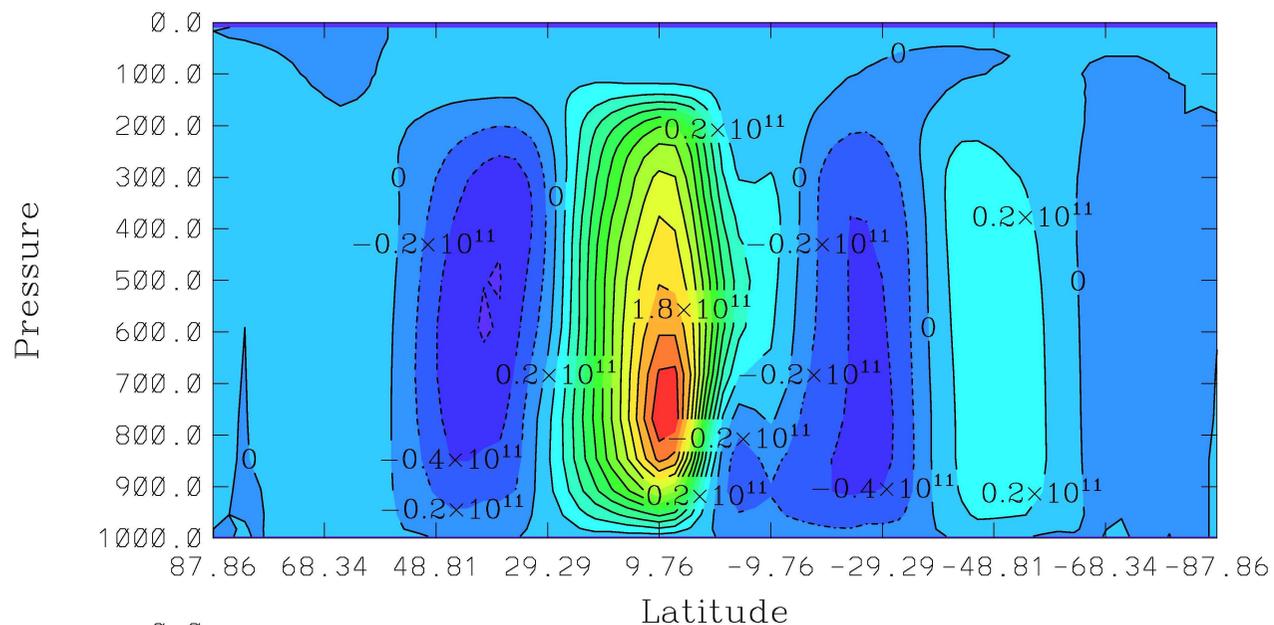
UM mean meridional circulation



Larkin (2000)

UM mean meridional circulation (tropical heating experiment)





UGCM
mmc

Δ mmc
due O₃
shift

Multiple regression of observational data

$$y(x,t) = \sum_i \beta_i(x) f_i(t) + \text{noise}$$

- $y(x,t)$ are data
- $f_i(t)$ is time-dependent climate factor i

10 factors are taken into account:

trend

solar irradiance

volcanic aerosol

ENSO

NAO

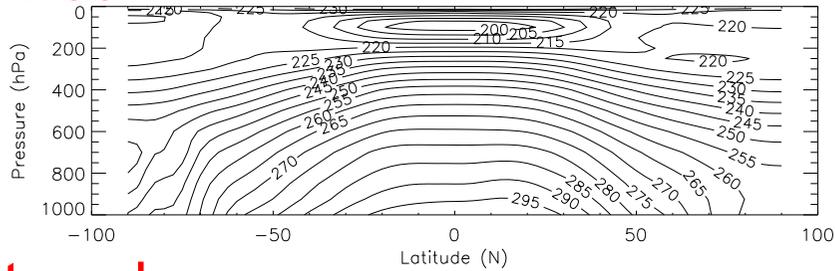
QBO

amplitude & phase of annual cycle

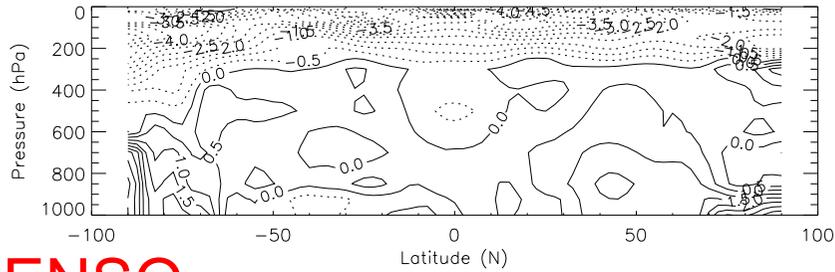
amplitude & phase of semi-annual cycle

Regression results: NCEP Tbar

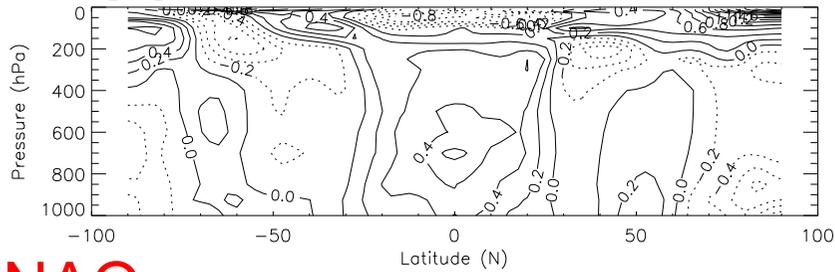
mean



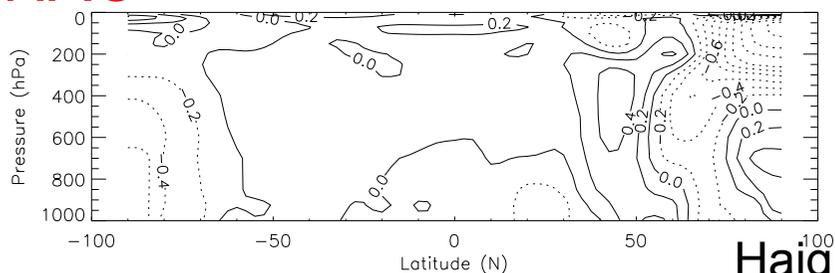
trend



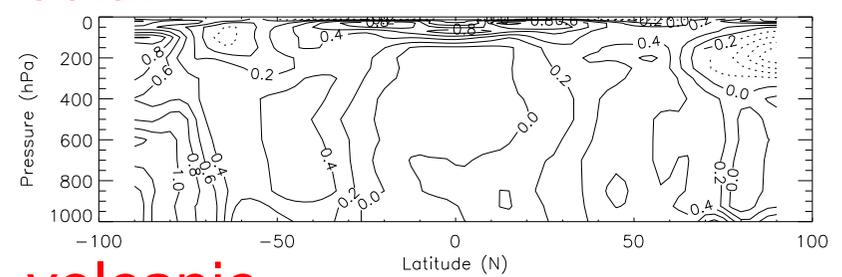
ENSO



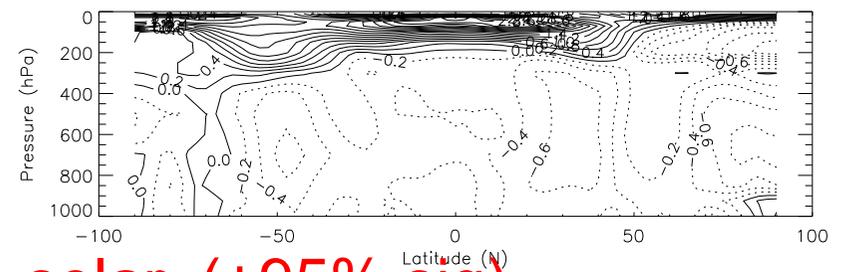
NAO



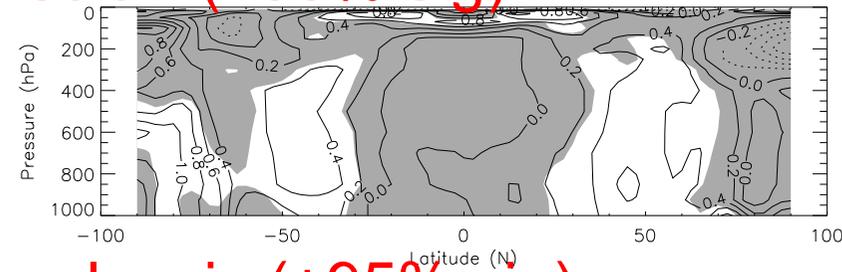
solar



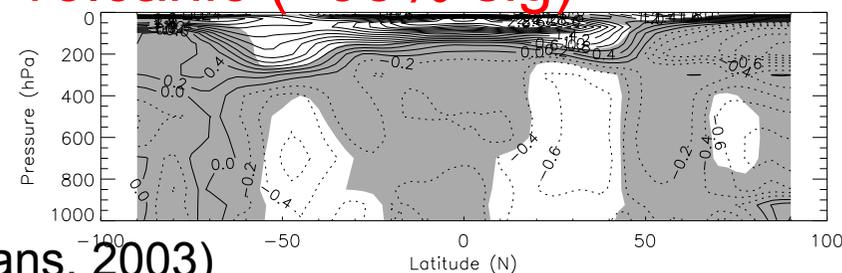
volcanic



solar (+95% sig)

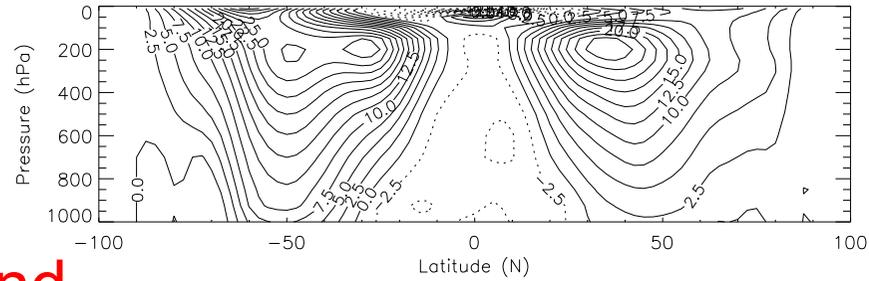


volcanic (+95% sig)

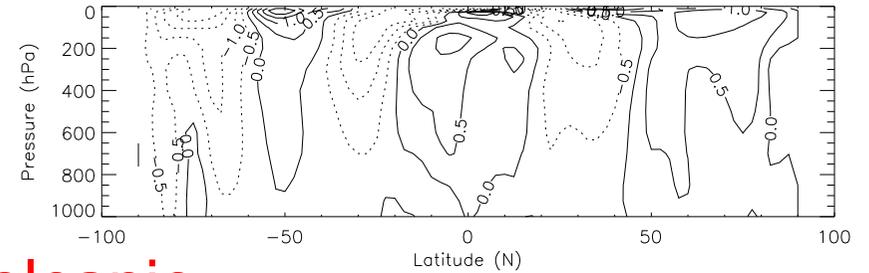


Regression results: NCEP ubar

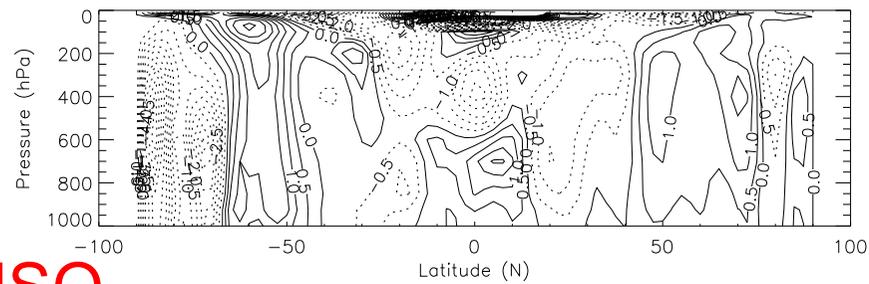
mean



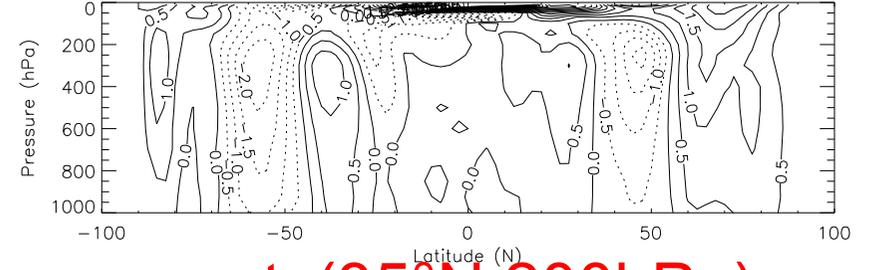
solar



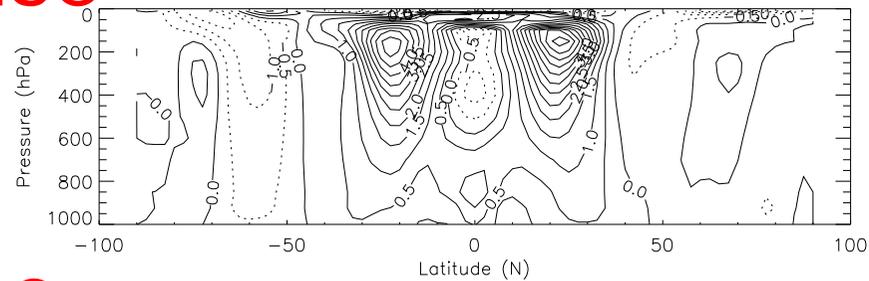
trend



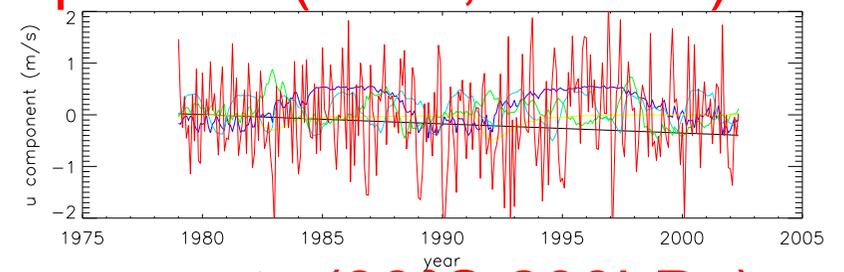
volcanic



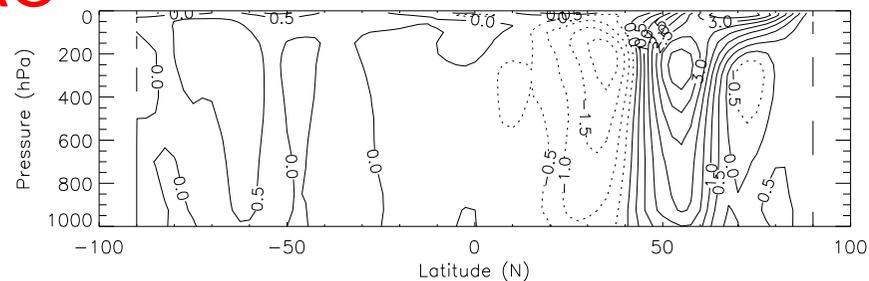
ENSO



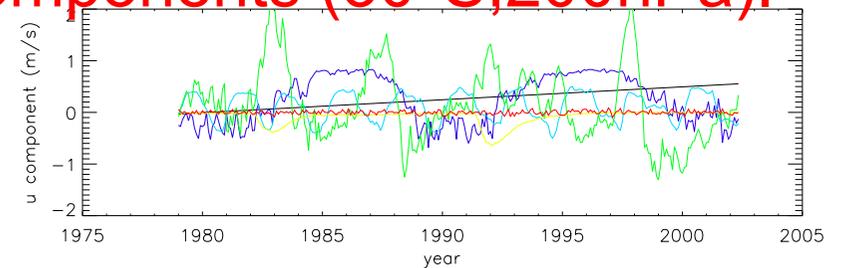
components(35°N,200hPa)



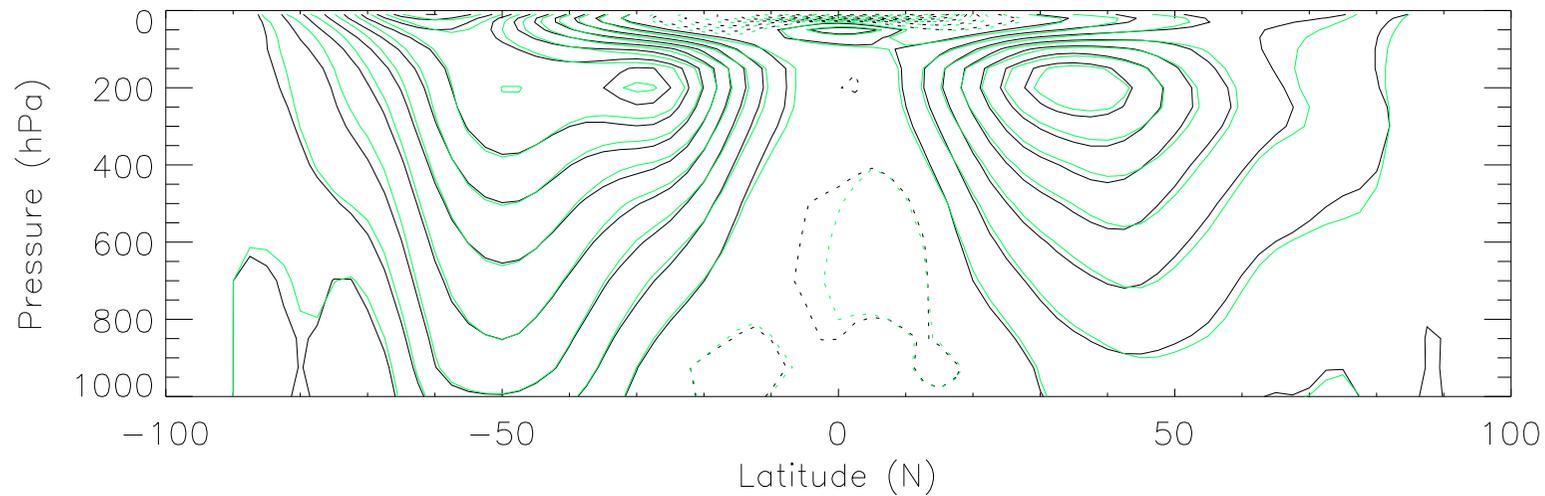
NAO



components (30°S,200hPa).

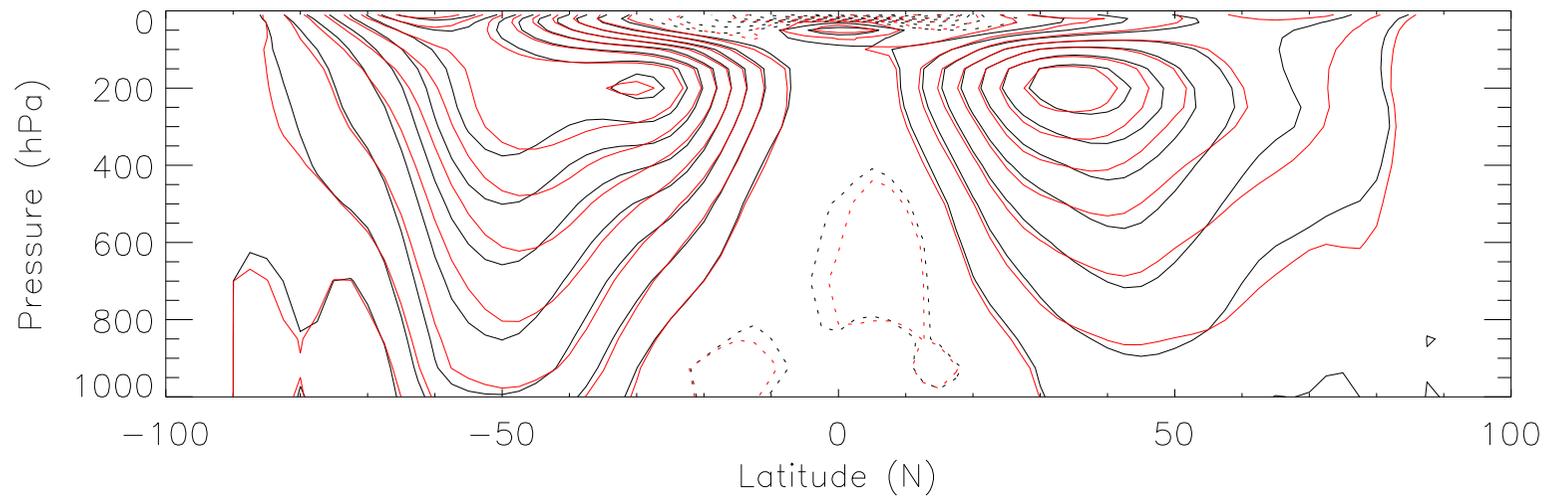


NCEP ubar regression results



solar min

solar max



low aerosol

PinaTubo

Dynamical core model experiments

As Held and Suarez (1994):

Full dynamics T42 L20. No orography.

Newtonian cooling (equinoctial radiative equilibrium temperatures). Rayleigh friction.

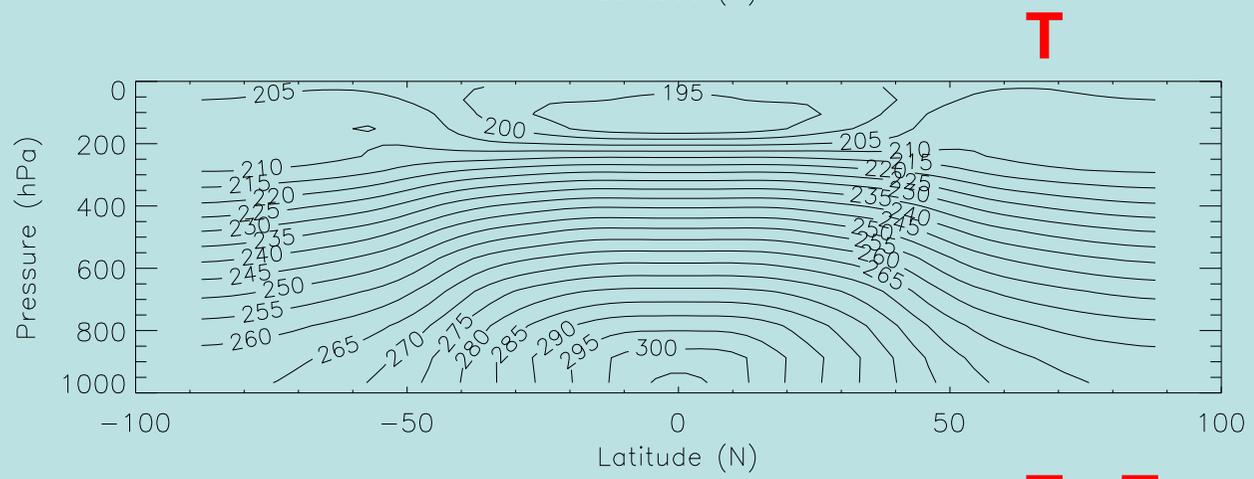
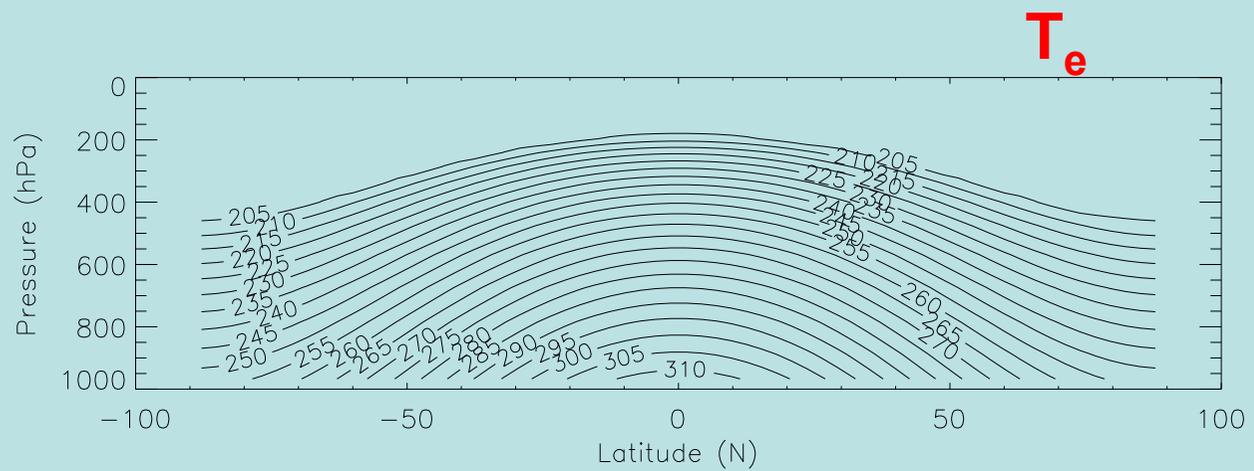
All experiments involved heating the lower stratosphere:

- 1K at all latitudes
- 1K at equator, $\cos^2(\text{lat})$ variation
- 5K at all latitudes
- 5K at equator, $\cos^2(\text{lat})$ variation

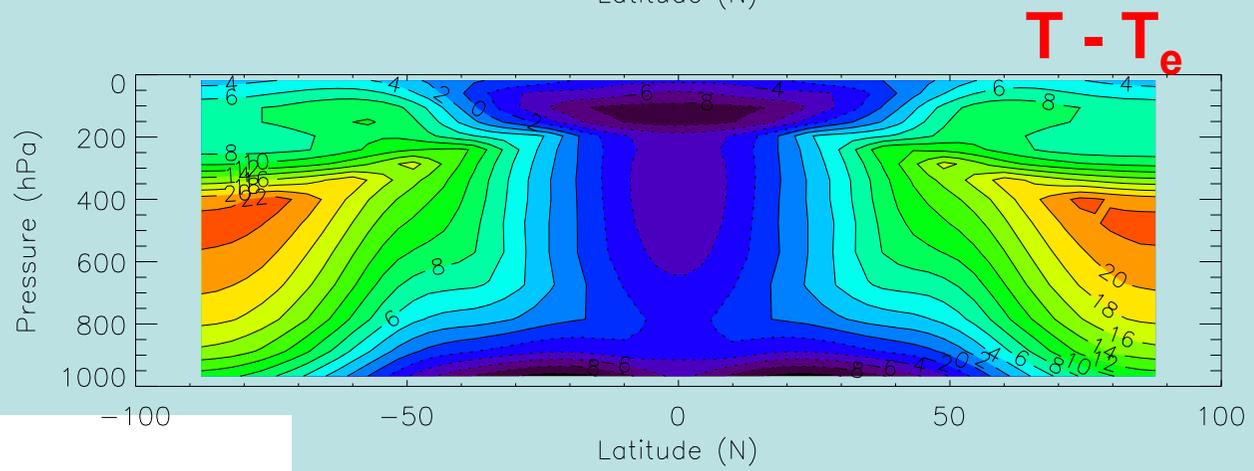
All runs 1200 days.

Temperatures

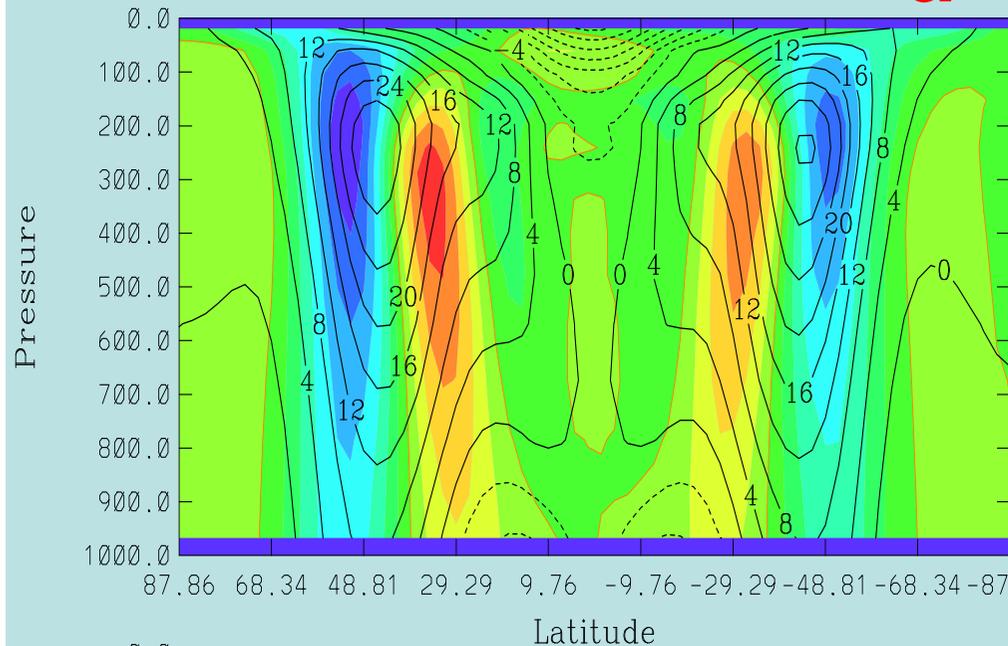
rad. eq.



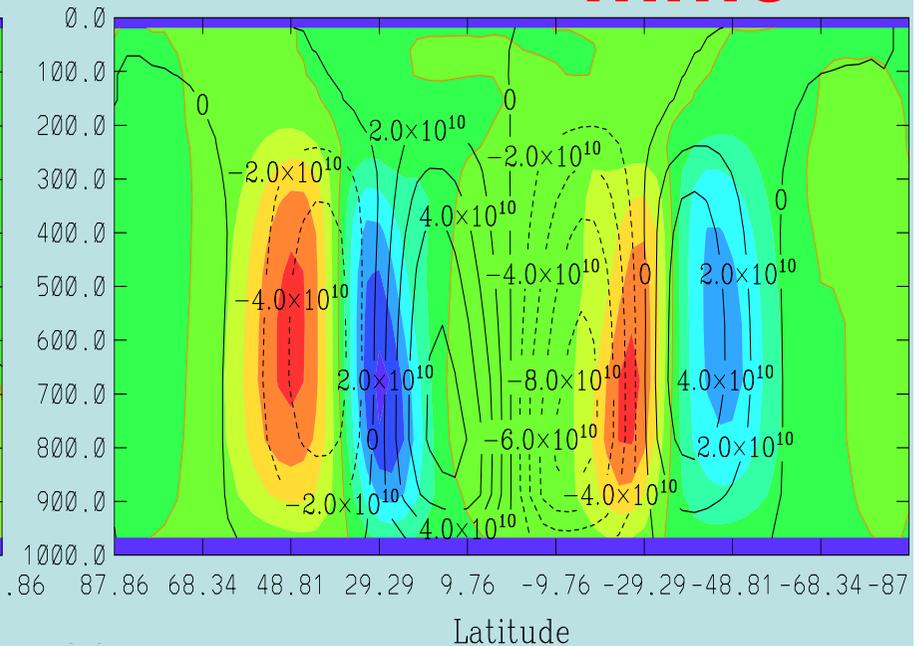
control run



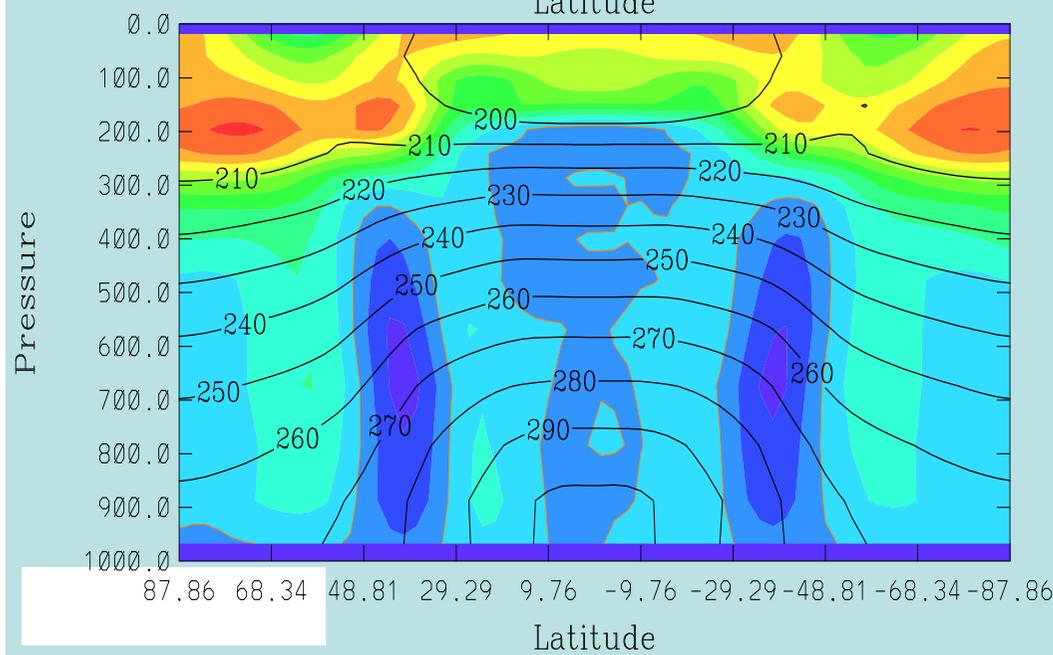
u



mmc

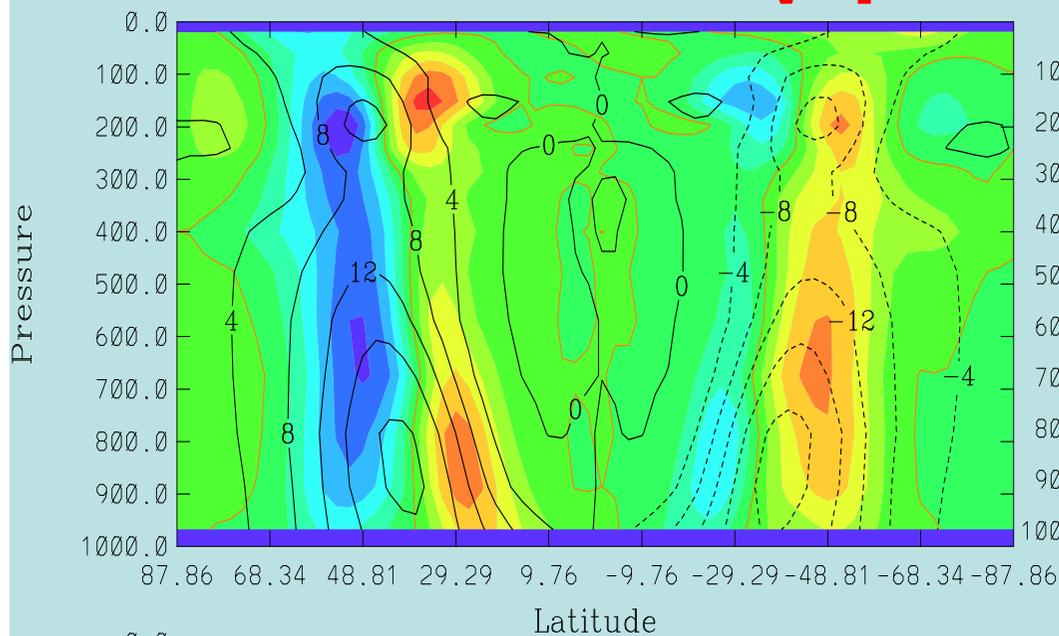


T

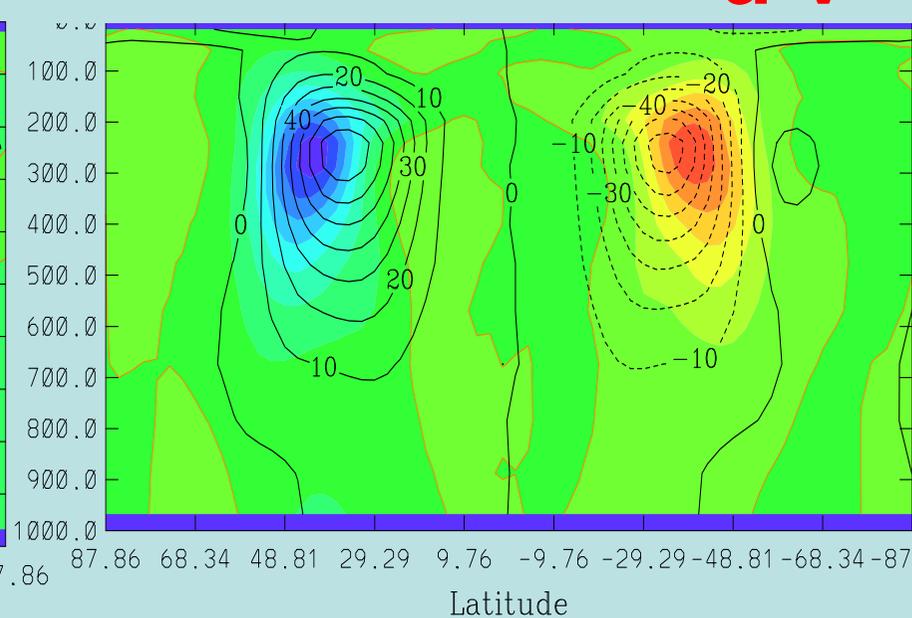


Uniform heating expt

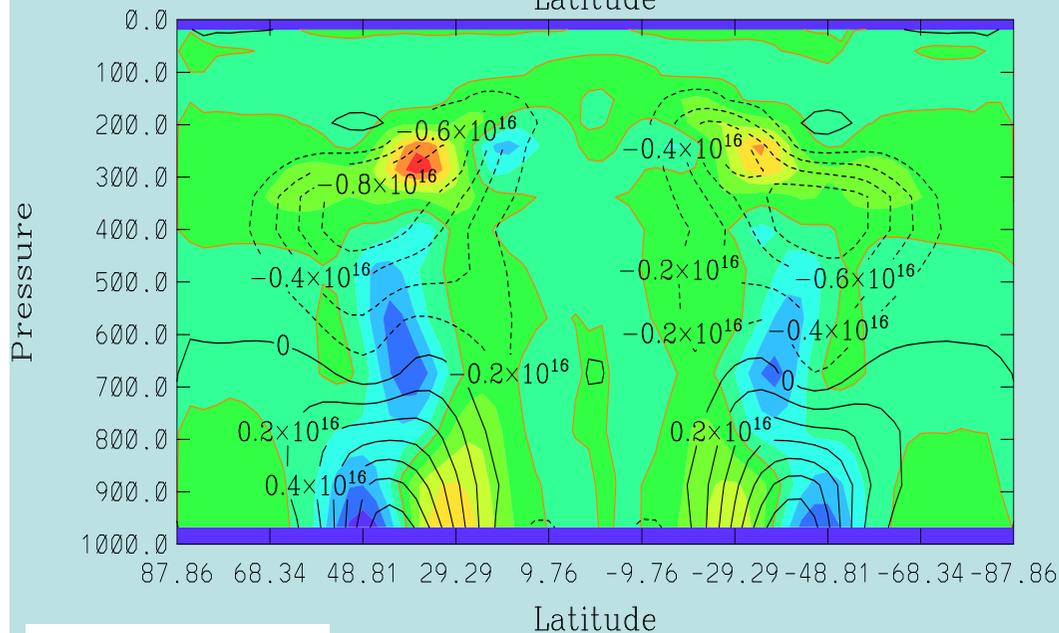
$v'T'$



$u'v'$

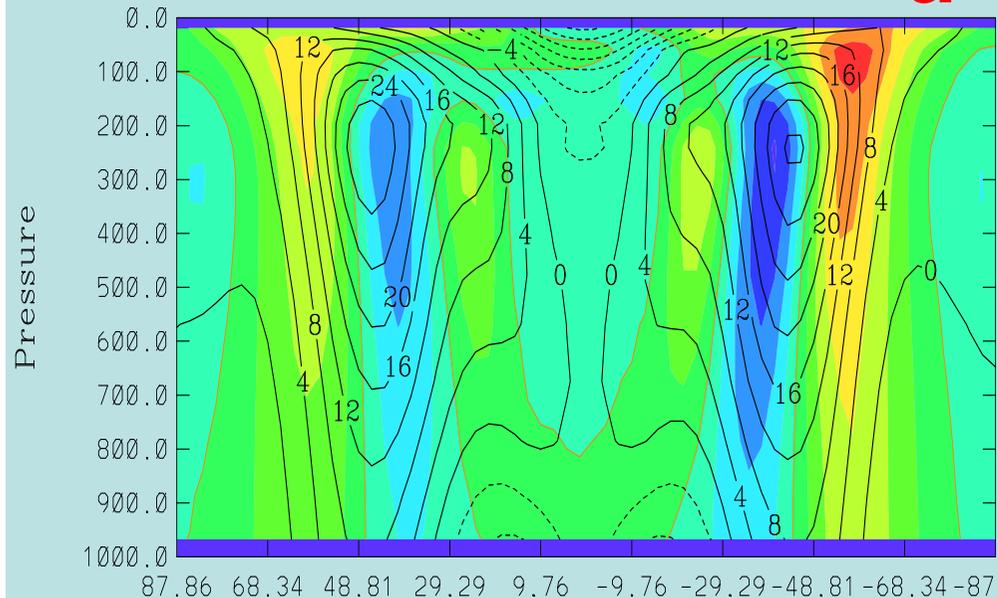


$\nabla \cdot F$

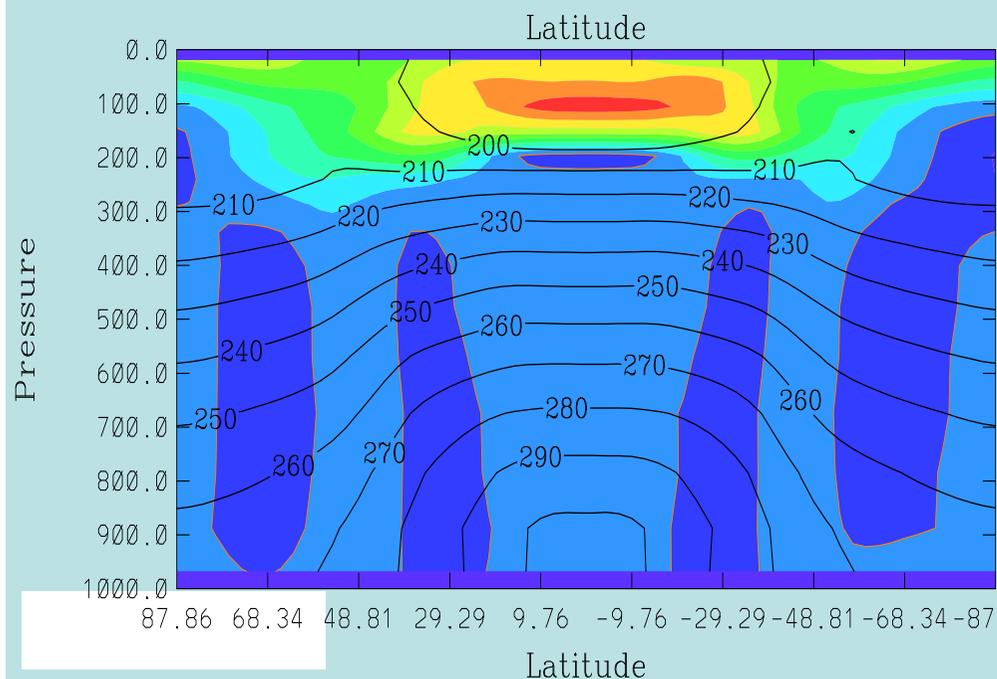
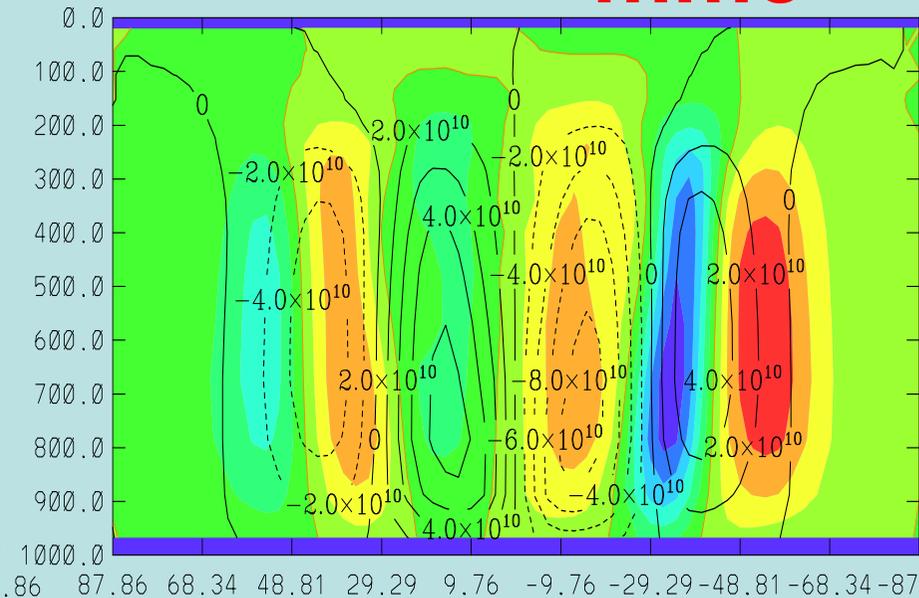


Uniform heating expt

u



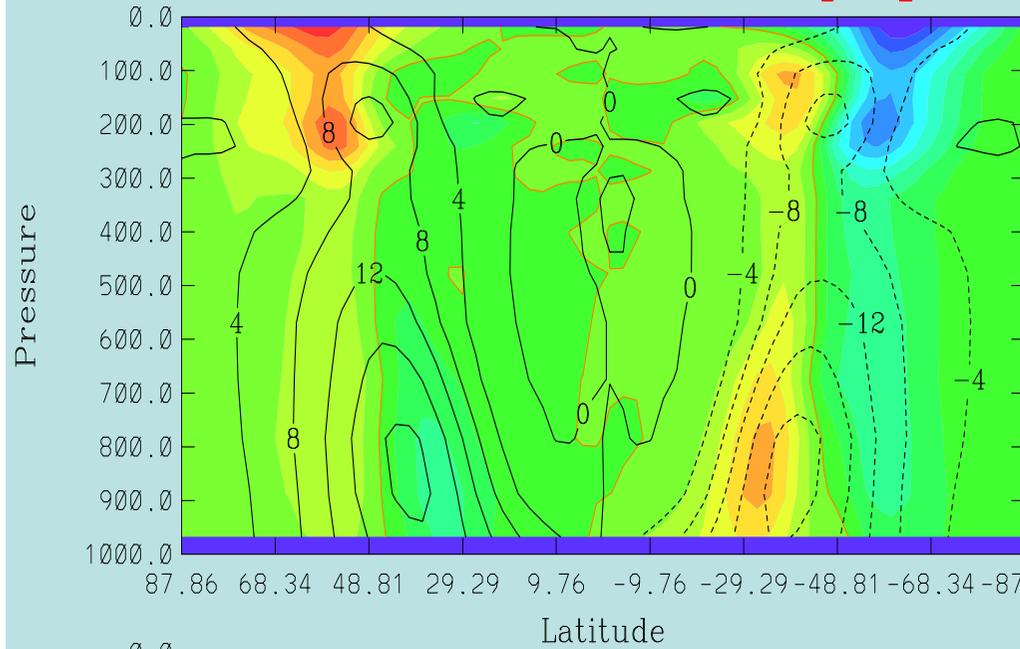
mmc



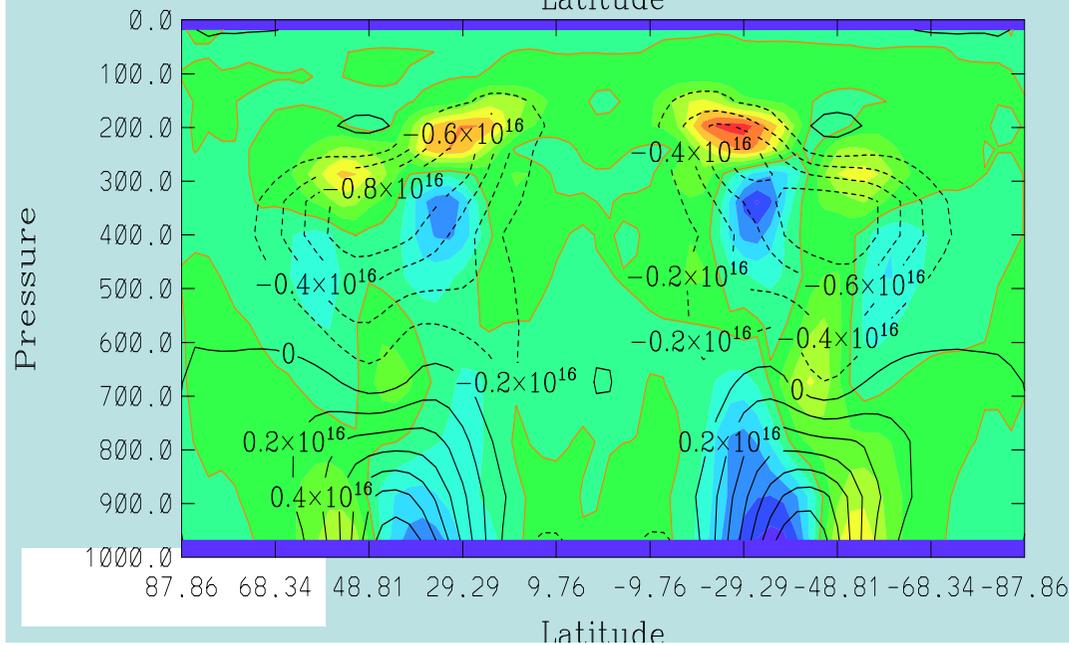
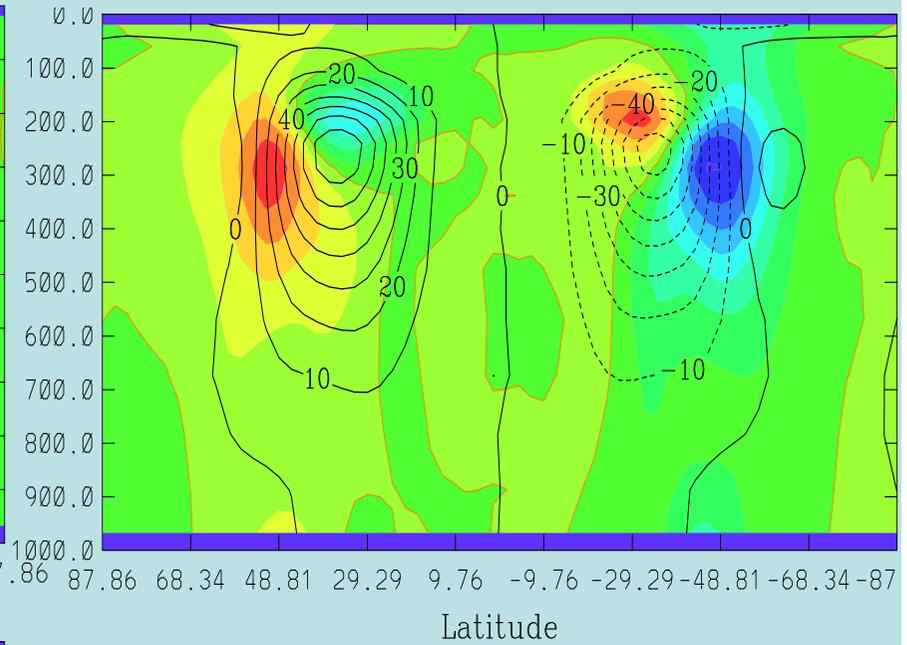
T

$\cos^2(\text{lat})$
heating expt.

$v'T'$



$u'v'$

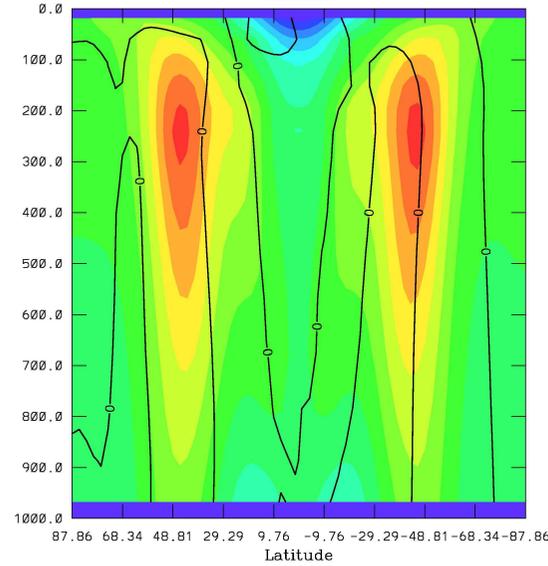
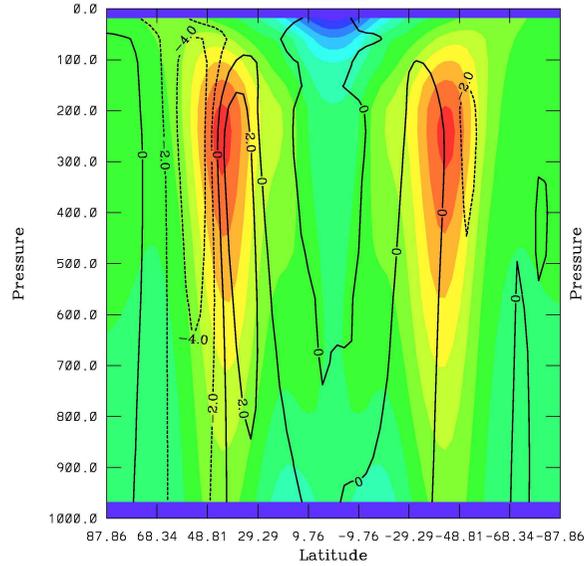
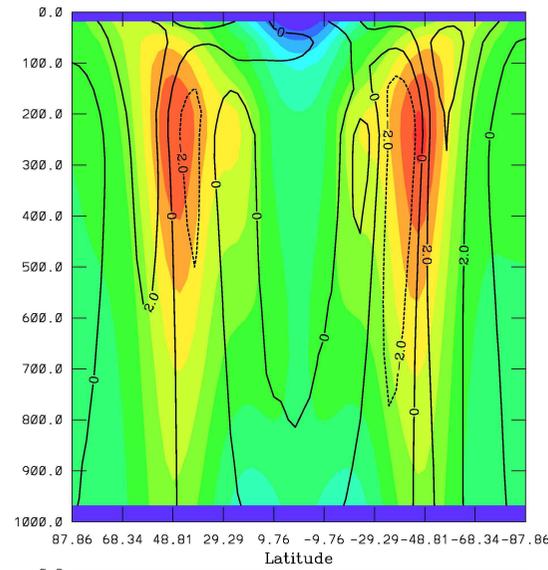
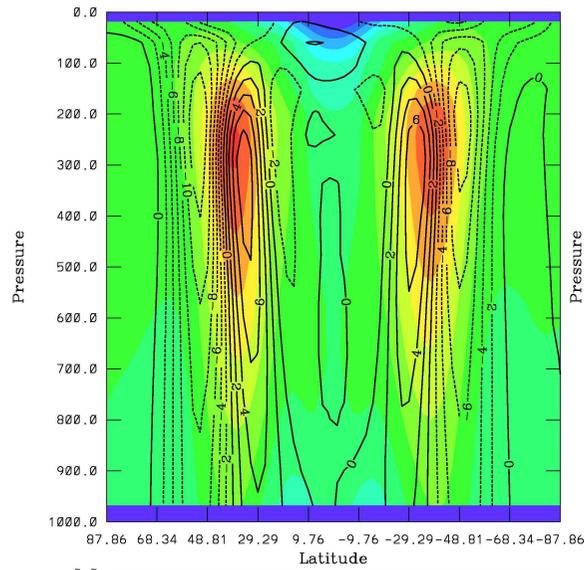


$\nabla \cdot \text{EPF}$

$\cos^2(\text{lat})$
heating expt.

uniform heating

$\cos^2(\text{lat})$ heating



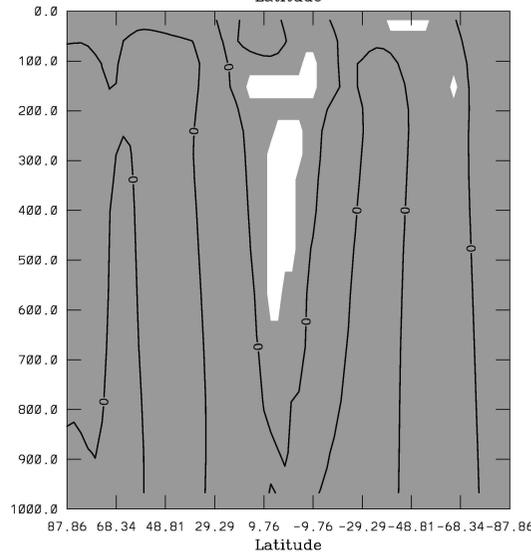
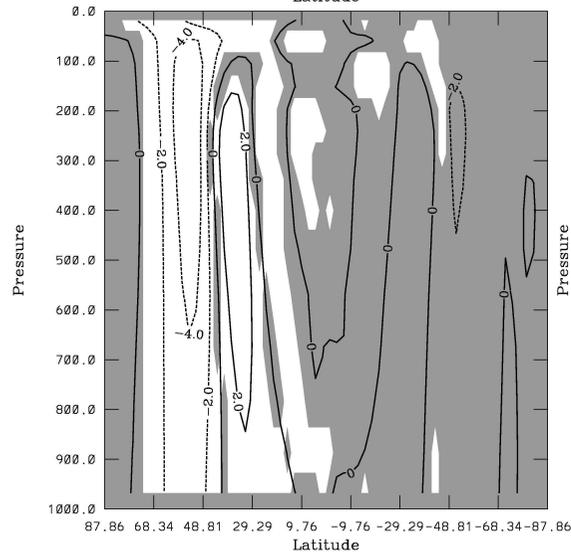
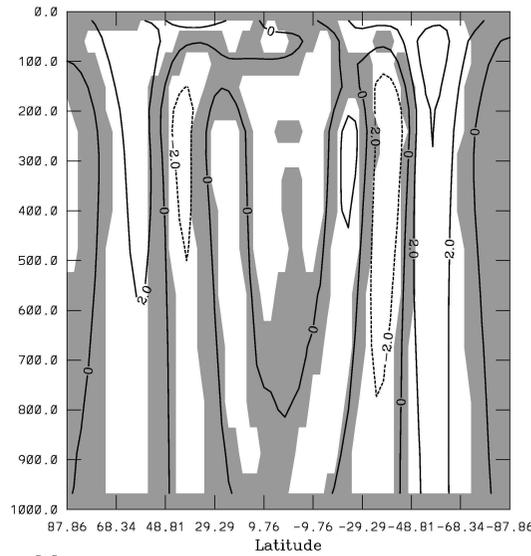
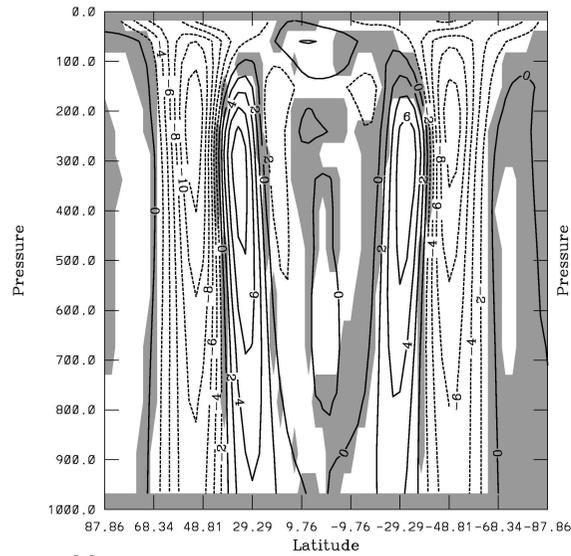
ubar

5K heating

1K heating

uniform heating

$\cos^2(\text{lat})$ heating



ubar

5K heating

shaded areas not statistically significant at 95% level

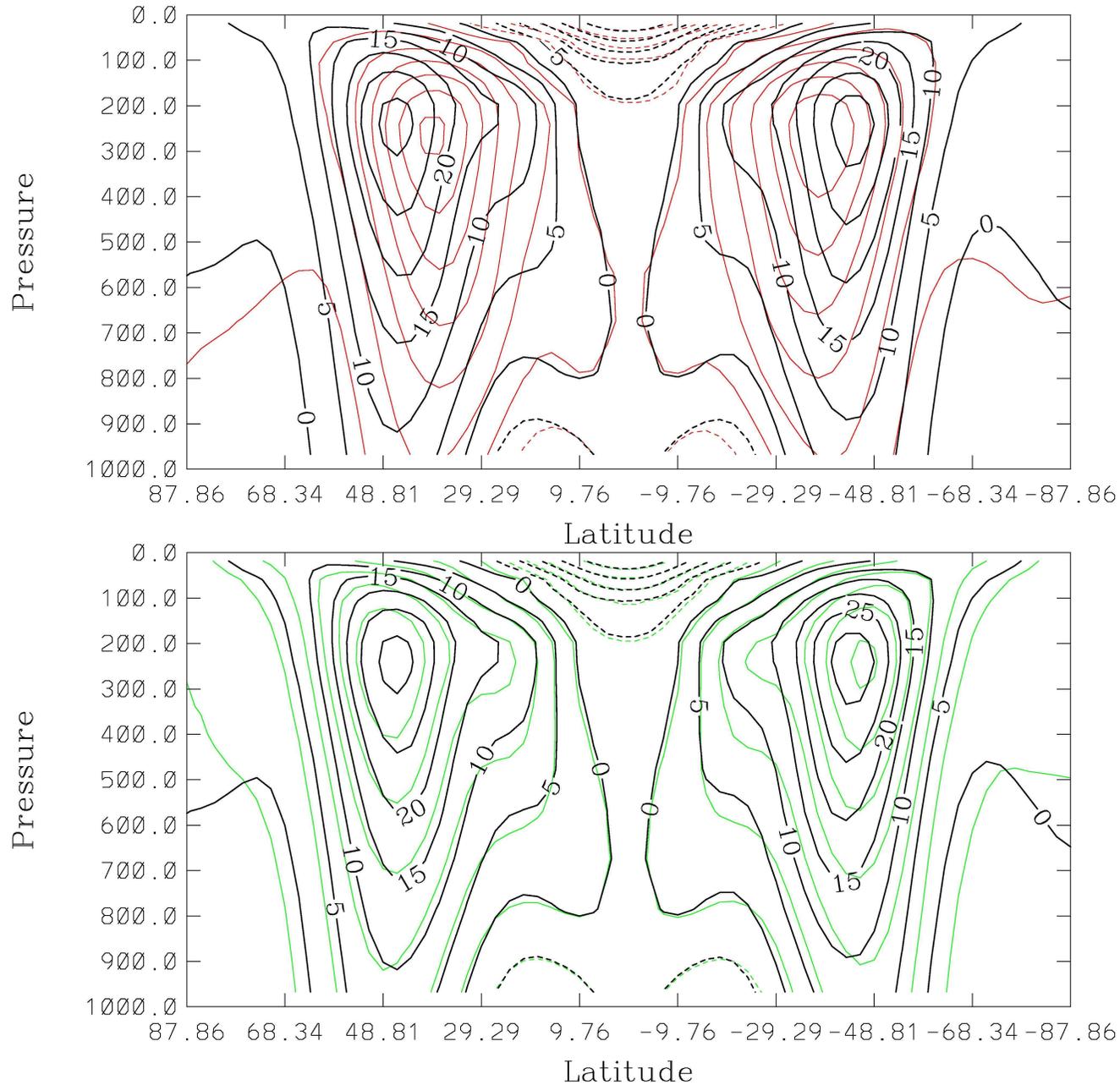
1K heating

ubar

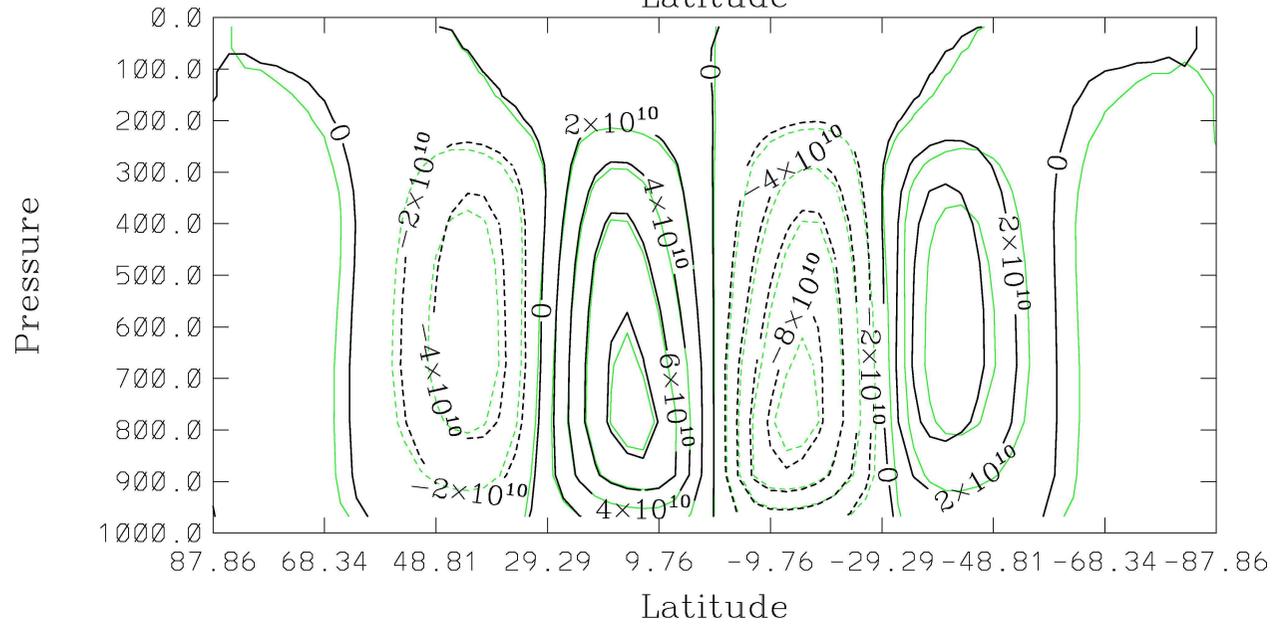
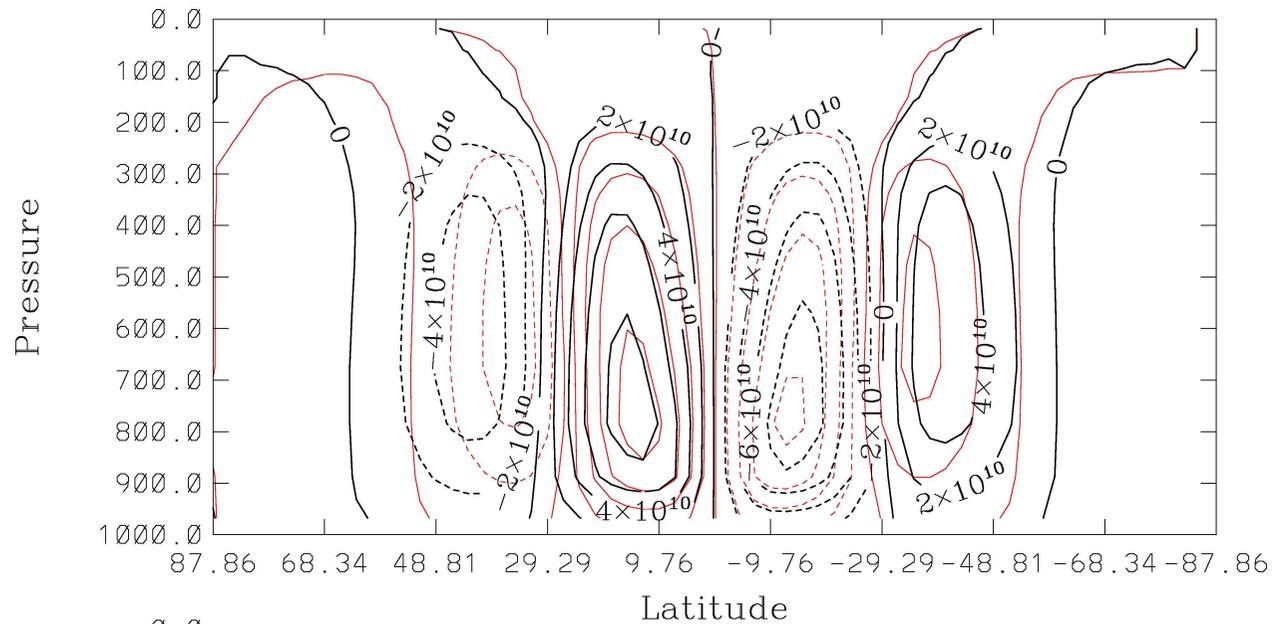
control

uniform heating

$\cos^2(\text{lat})$ heating



mean meridional circulation

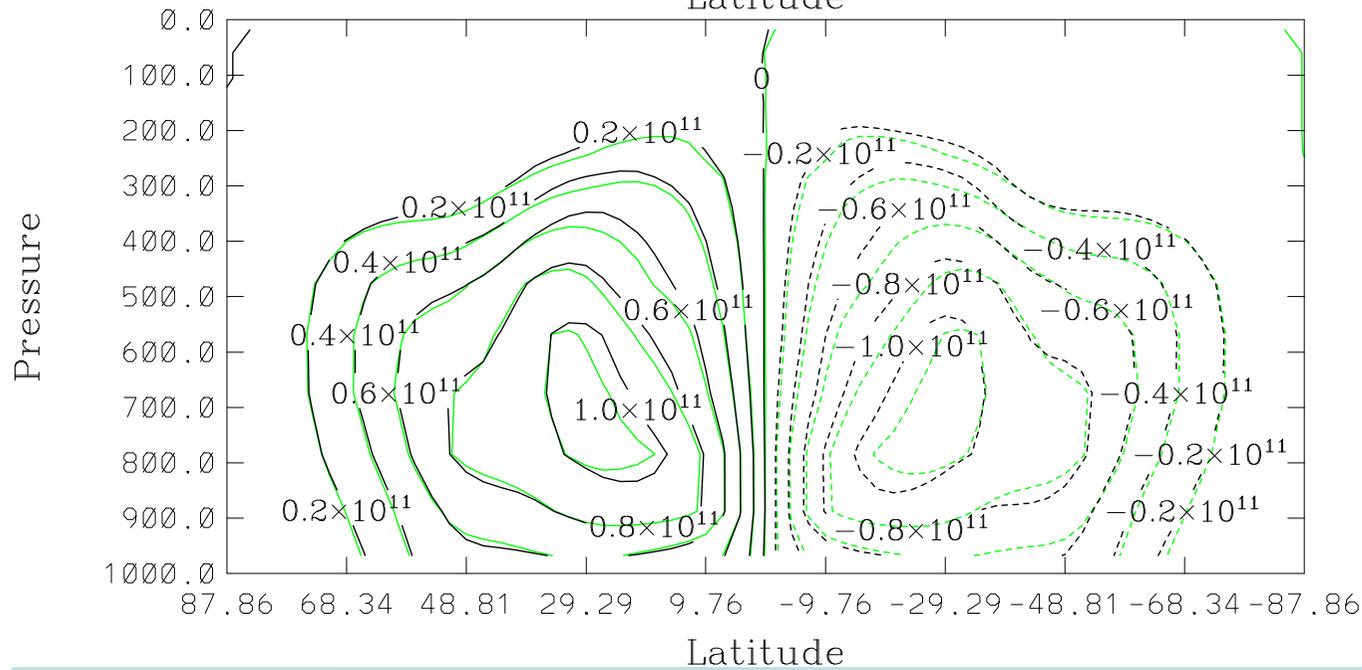
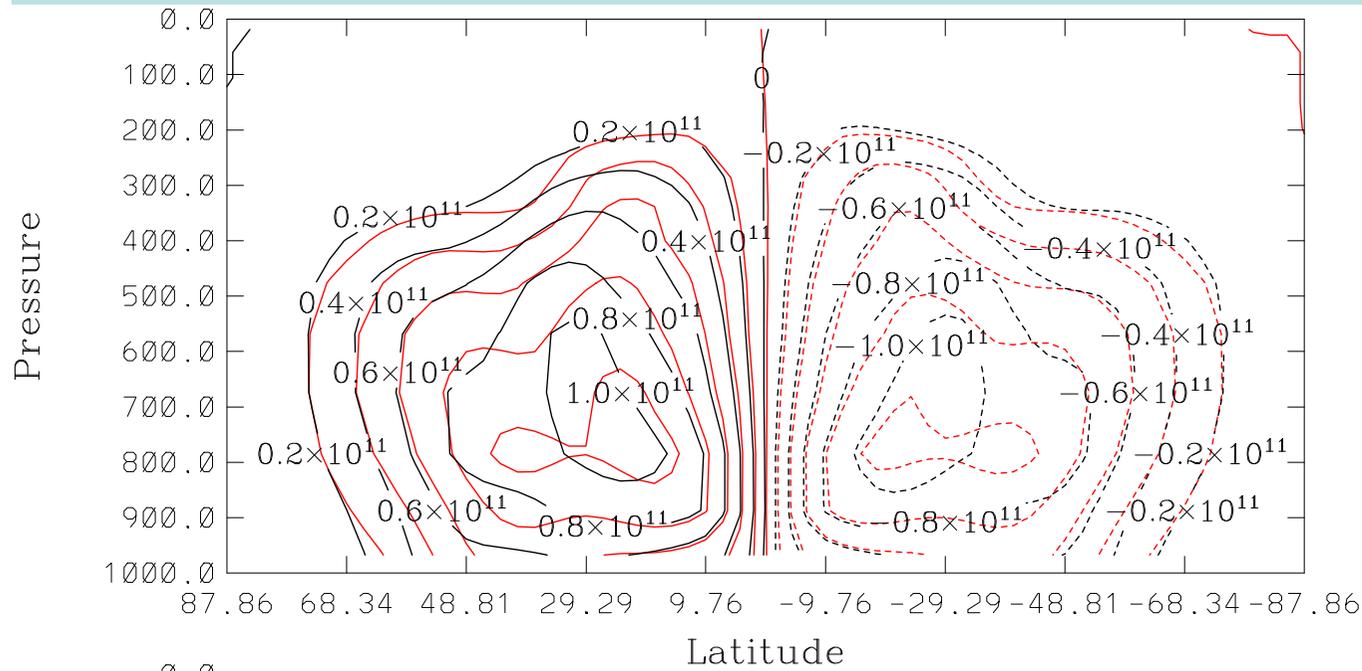


control

uniform heating

$\cos^2(\text{lat})$ heating

TEM residual circulation



control

uniform heating

cos²(lat) heating

Conclusions - I

- Model simulations show a characteristic pattern of response to enhanced solar UV: the sub-tropical jets weaken & move poleward, the Hadley cells weaken.
- The same patterns, with similar magnitudes, are found in multiple regression studies of observational data.

Conclusions - II

- Results from a simplified model suggest that:
 - (a) Heating of the lower stratosphere tends to weaken the sub-tropical jets as a result of reduced eddy activity.
 - (b) The distribution of the heating in the stratosphere determines any shift in position of the jets.